



COMPASS

---

---

# ***Particle identification using*** ***Ring Cherenkov detector*** ***technology at COMPASS***

**Miroslav Finger**

*Charles University in Prague, Czech Republic*  
*Faculty of Mathematics and Physics*

- 1. COMPASS Experiment at SPS CERN**
- 2. RIing Imaging CHerenkov Detectors Basics**
- 3. COMPASS RICH-1 Detector for Particle Identification 2009/2010**
- 4. Conclusions**

# Spin observables → experiments



@RHIC:

PHENIX

STAR

BRAHMS

NICA&NUCLOTRON @JINR



hermes @DESY



@KEK



@CERN

Jefferson Lab  
@6GeV

*far future*

PANDA&PAX@GSI JLab@12GeV EIC JParc IHEP

future TMD  
measurements

pp

sidis

e<sup>+</sup>e<sup>-</sup>

RHIC PAX

HERMES COMPASS

JLab

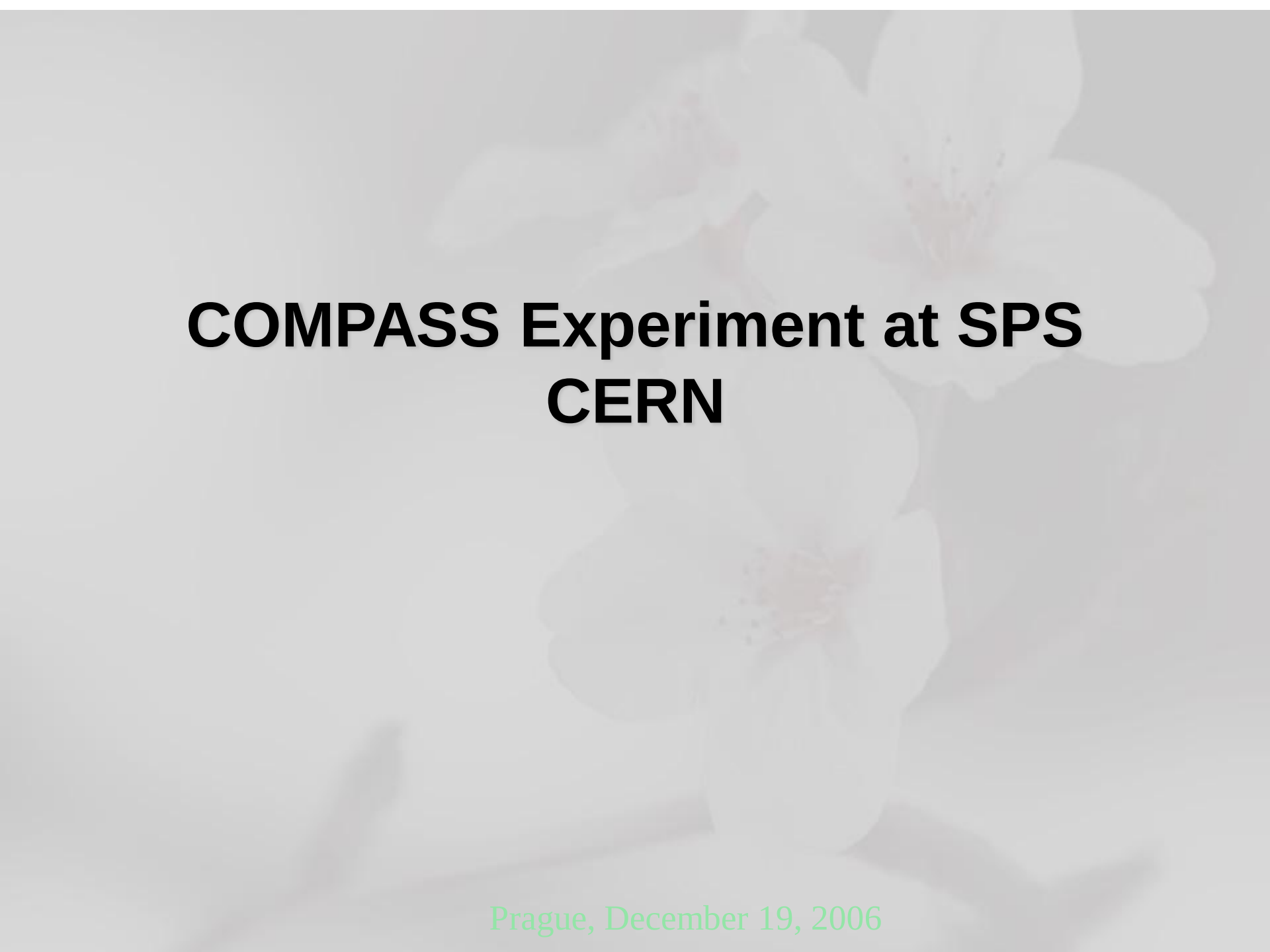
NUCLOTRON

JParc

EIC

BELLE

summary

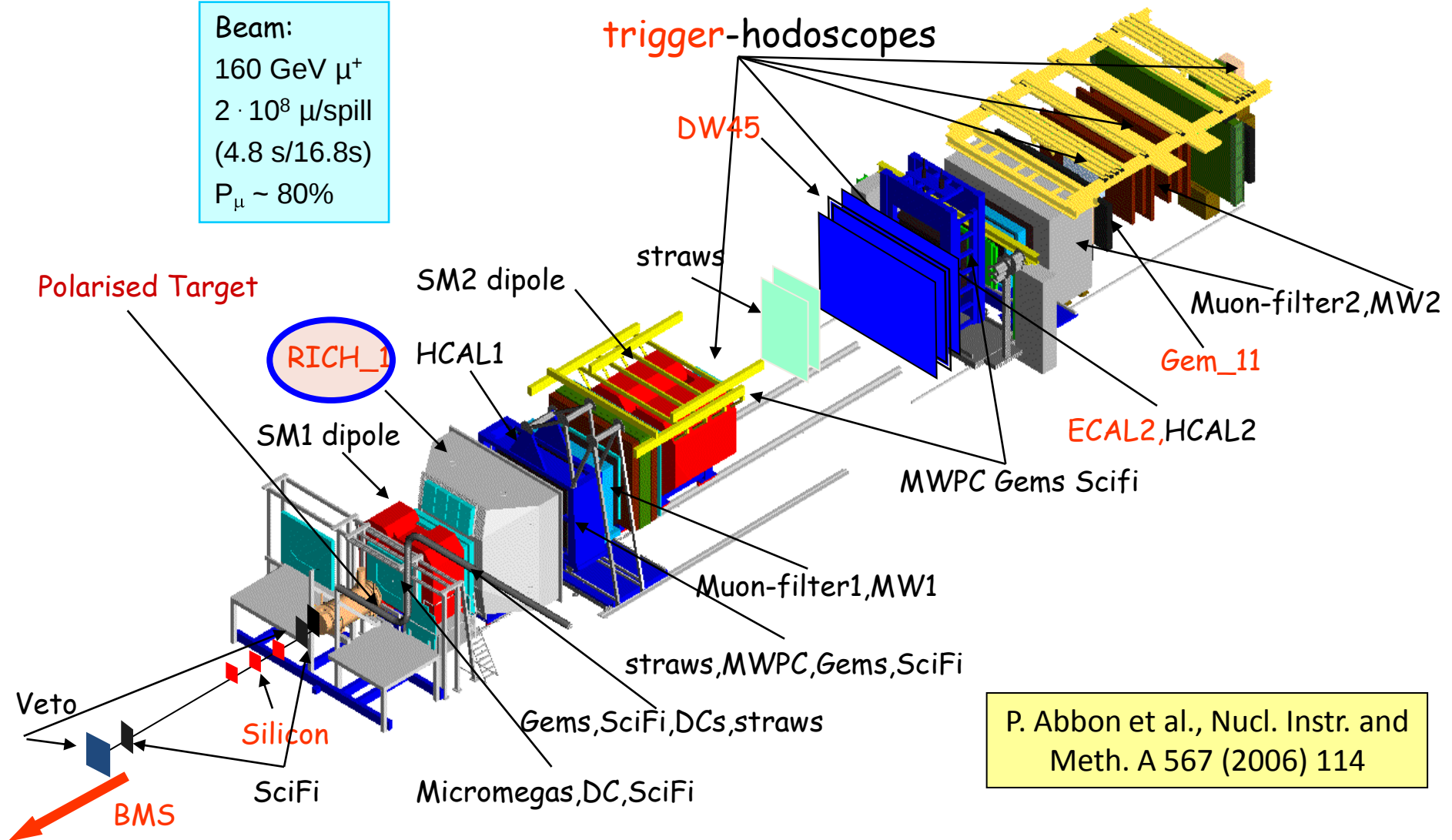


# **COMPASS Experiment at SPS CERN**

Prague, December 19, 2006

# THE COMPASS SPECTROMETER

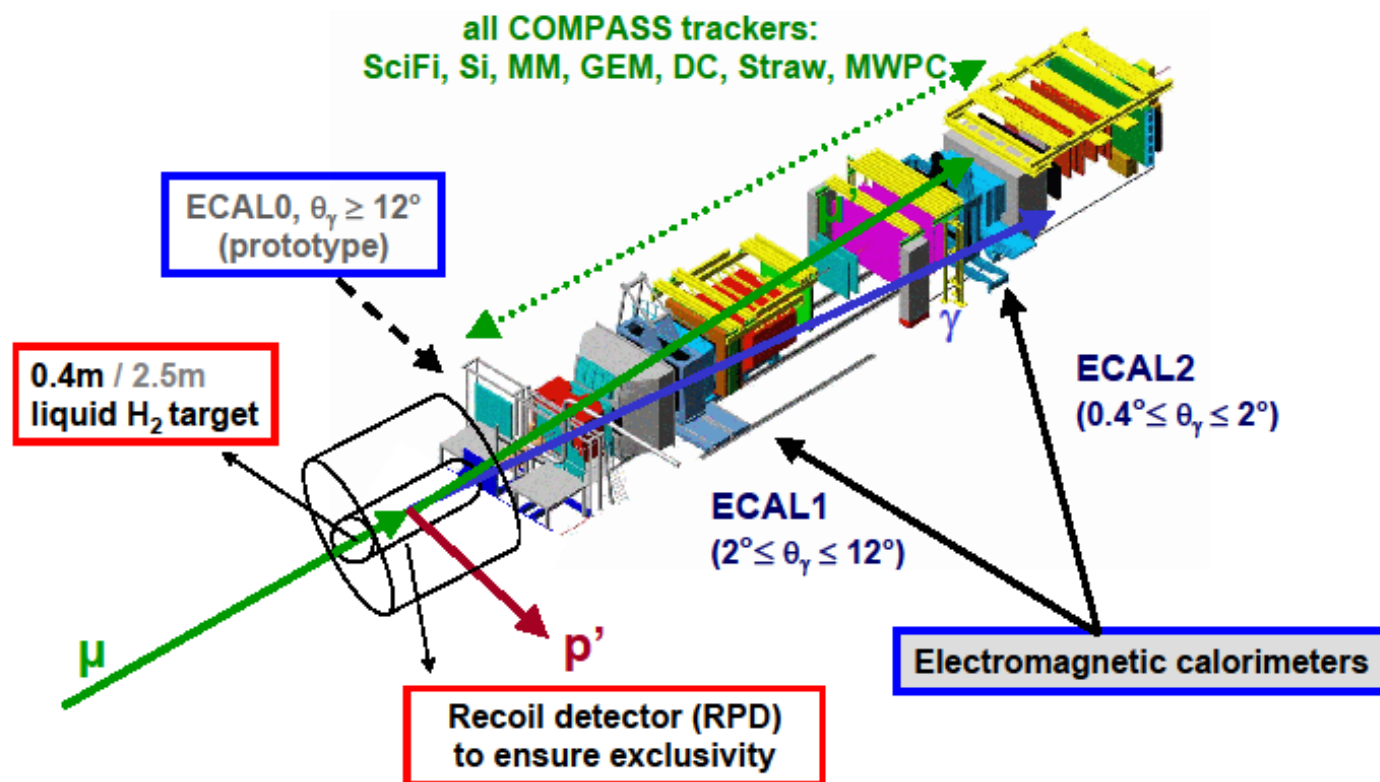
Beam:  
160 GeV  $\mu^+$   
 $2 \cdot 10^8$   $\mu$ /spill  
(4.8 s/16.8s)  
 $P_\mu \sim 80\%$



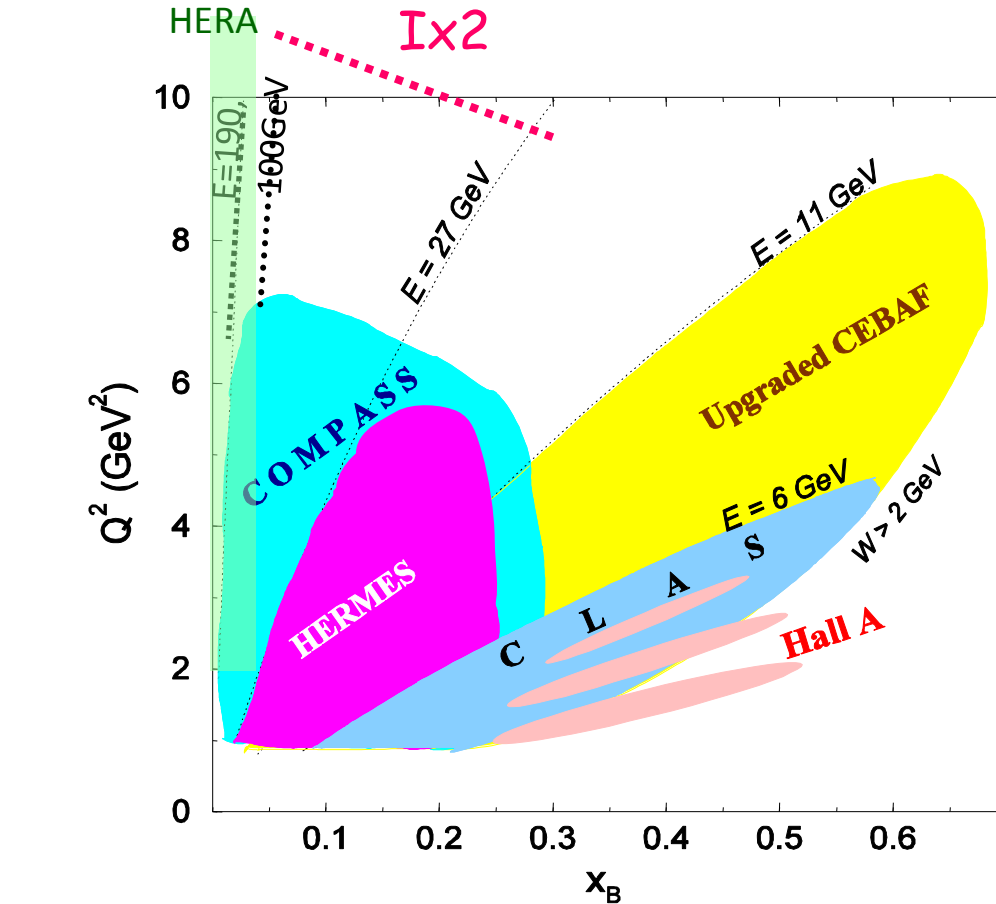
P. Abbon et al., Nucl. Instr. and Meth. A 567 (2006) 114



## COMPASS: Setup 2008 and beyond



# Competition in the world and COMPASS role



Gluons  
 valence quarks and sea quarks and gluons  
 COMPASS 2010  
 valence quarks  
 JLab 12 GeV, FAIR, ... 2014

## COMPASS at CERN-SPS

High energy muon beam  
100/190 GeV

$\mu^+$  or  $\mu^-$   
 change once per day  
 $\text{polar}(\mu^+) = -0.80$   
 $\text{polar}(\mu^-) = +0.80$

$2 \cdot 10^8 \mu$  per SPS cycle

if intensity  $\times 2$ ,  
 $Q^2$  range up to 11 GeV<sup>2</sup>

after 2010,  
which improvements  
on the muon line ?

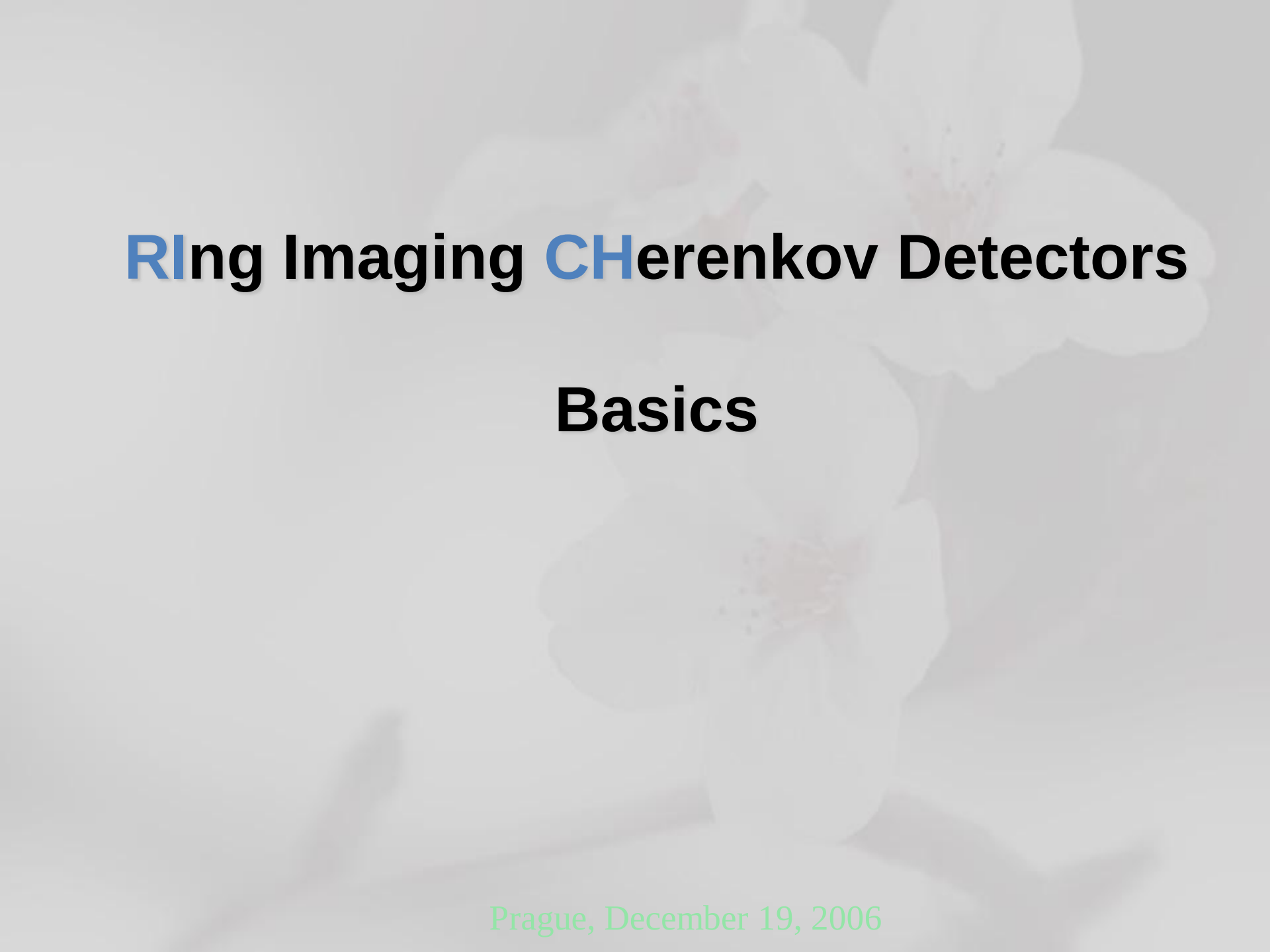
# Physics program of COMPASS - 1

- Experiments with muon beam

- $\Delta G/G$
- $g_1$
- Transverse spin effects
- Flavor decomposition of spin distribution functions
- Vector meson production
- Spin transfer in  $\Lambda$ -hyperon production

- Experiments with hadron beams

- Pion and Kaon polarizabilities
- Diffractive production of exotic states
- Search for glueballs
- Light meson spectroscopy
- Production of double charmed baryons

The background of the slide features a faint, artistic image of white flowers, possibly cherry blossoms, with visible petals and stamens. The image is light and serves as a subtle backdrop for the text.

# **RI**ng Imaging **CH**erenkov Detectors

## **Basics**

Prague, December 19, 2006

A charged particle with a velocity  $v$  *larger than* the velocity of light in a medium emits light.

Pavel A. Cherenkov, Ilja M. Frank, Igor Y. Tamm  
Nobel Price 1958



## Basic Formulas

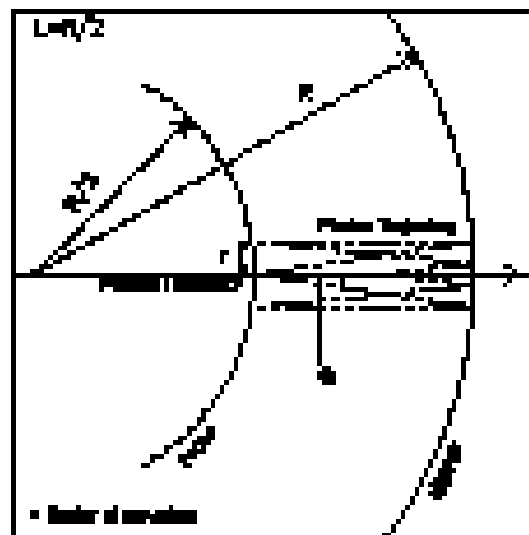
Threshold:  $\beta_{\text{thres}} = \frac{V_{\text{thres}}}{c} = \frac{1}{n} \quad \gamma_{\text{thres}} = \frac{n}{\sqrt{n^2 - 1}}$

Angle of emission:  $\cos \theta_c = \frac{1}{\beta n} = \frac{1}{\frac{v}{c} n}$

Number of photons:  $\frac{d^2 N}{dE dl} = \frac{\alpha z^2}{\hbar c} \left( 1 - \frac{1}{(\beta n)^2} \right) = \frac{\alpha z^2}{\hbar c} \sin^2 \theta_c$   
 $\frac{d^2 N}{d\lambda dl} = \frac{2\pi \alpha z^2}{\lambda^2} \sin^2 \theta_c$

# Ring Imaging Cherenkov – The Basics

$$\cos \theta_c = \frac{1}{\beta \cdot n} \quad r = F \cdot \tan \theta_c = \frac{R}{2} \cdot \tan \theta_c \quad N_{ph} = N_0 \cdot L \cdot \sin^2 \theta_c$$



$\theta_c$ : Cherenkov angle

$\beta$ : velocity

$n$ : refractive index

$r$ : Radius of ring on focal surface

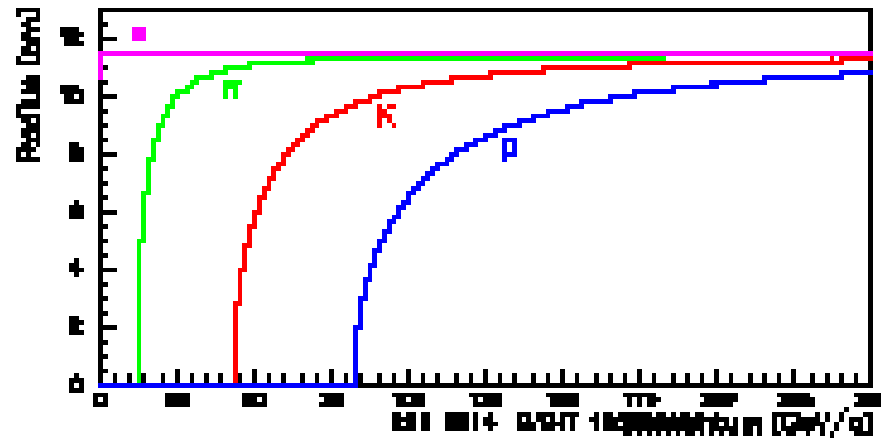
$R$ : Radius of curvature of  
spherical mirror(s)

$F$ : Focal length ( $F = R/2$ )

$L$ : Radiator length (usually  $L = F$ )

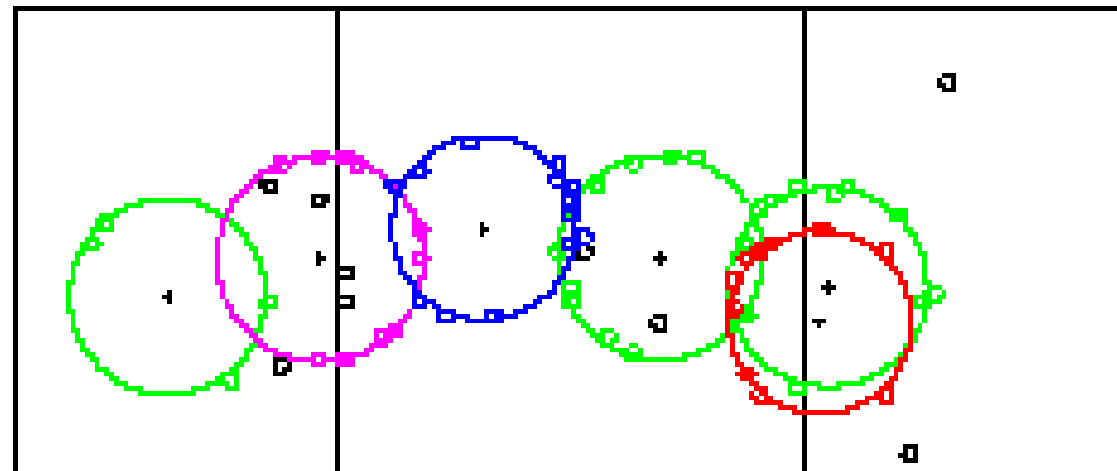
Parallel particles have the same ring image

Cherenkov Rod<sub>II</sub> – Neon Rodulator,  $F=1000\text{cm}$



small angle  
approximation:

$$r = \frac{R}{2} \sqrt{2 - \frac{2}{n} \sqrt{1 + \frac{m^2 c^2}{p^2}}}$$



- $$r = F \tan \Theta_c \approx \frac{R}{2} \sqrt{2 - \frac{2}{n} \sqrt{1 + \frac{m^2 c^2}{p^2}}}$$

Knowing  $R$  and  $n$ , and measuring  $r$ , we can identify a particle of mass  $m$ .

$$\Theta_c \Delta \Theta_c = \frac{m_1^2 - m_2^2}{2p^2}$$

Usually (production experiments):

Most difficult to separate pions and kaons

$\mu$  and  $e$  are identified by other means

But really: Most difficult is  $\mu$ - $\pi$ , key features of  
Velocity Spectrometer (CKM, NA62)

A RICH Detector is as simple as

- a box
- some mirrors
- and a few phototubes?

NO!

## Short History of RICHes in Experiments

**First Generation:** Beginning of 1980's.

Examples: E605, Omega (WA69, WA82), E665.

**Second Generation:** End-of 80's beginning of 90's.

Examples: Upgraded Omega (WA89, WA94),  
Delphi, SLD-GRID, CERES, SPHINX.

**Third Generation:** Mid-End 90's.

Examples: SELEX, Hermes, Hera-B.

**Fourth Generation:** BaBar-DIRC, PHENIX, CLEO-III,  
COMPASS, ...

**New Generation:** LHCb, ALICE, *BTeV*, *CKM*, ...

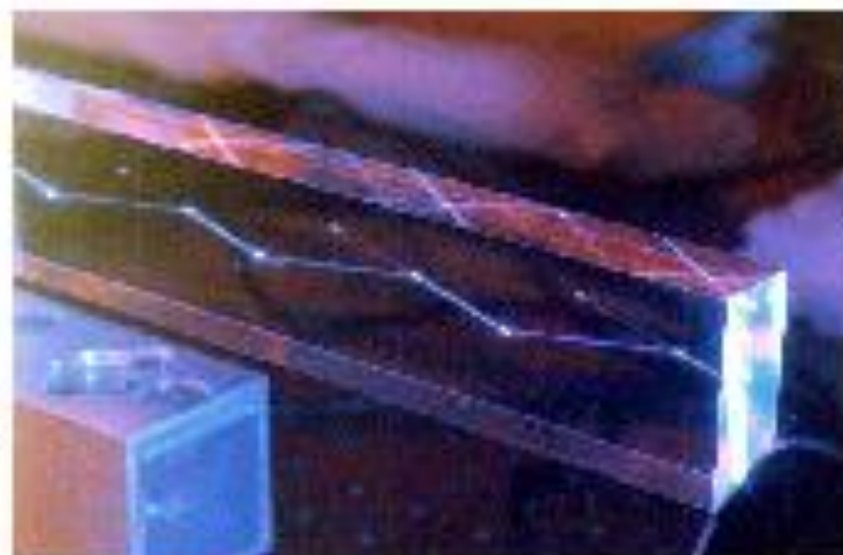
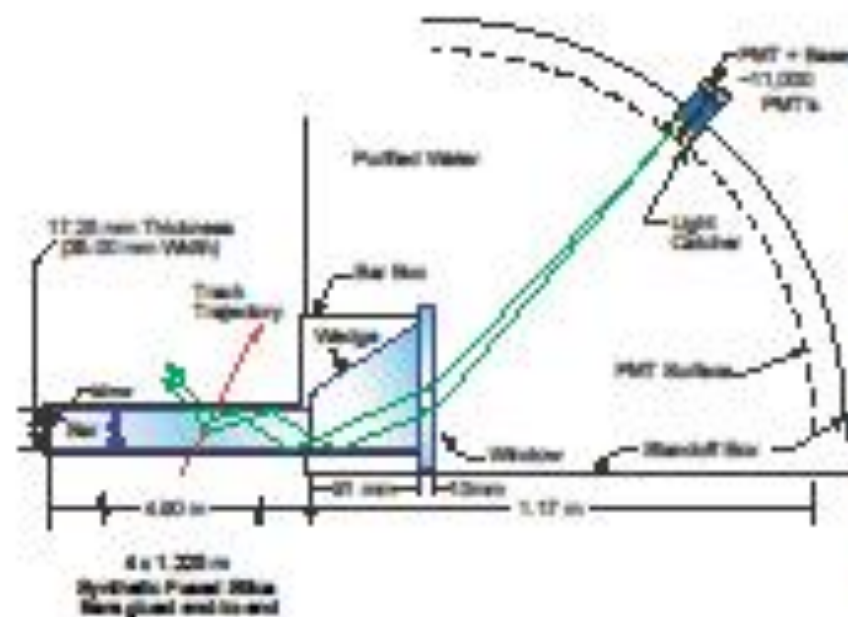
**Future:** NA62, Panda, WASA, CBM, ...

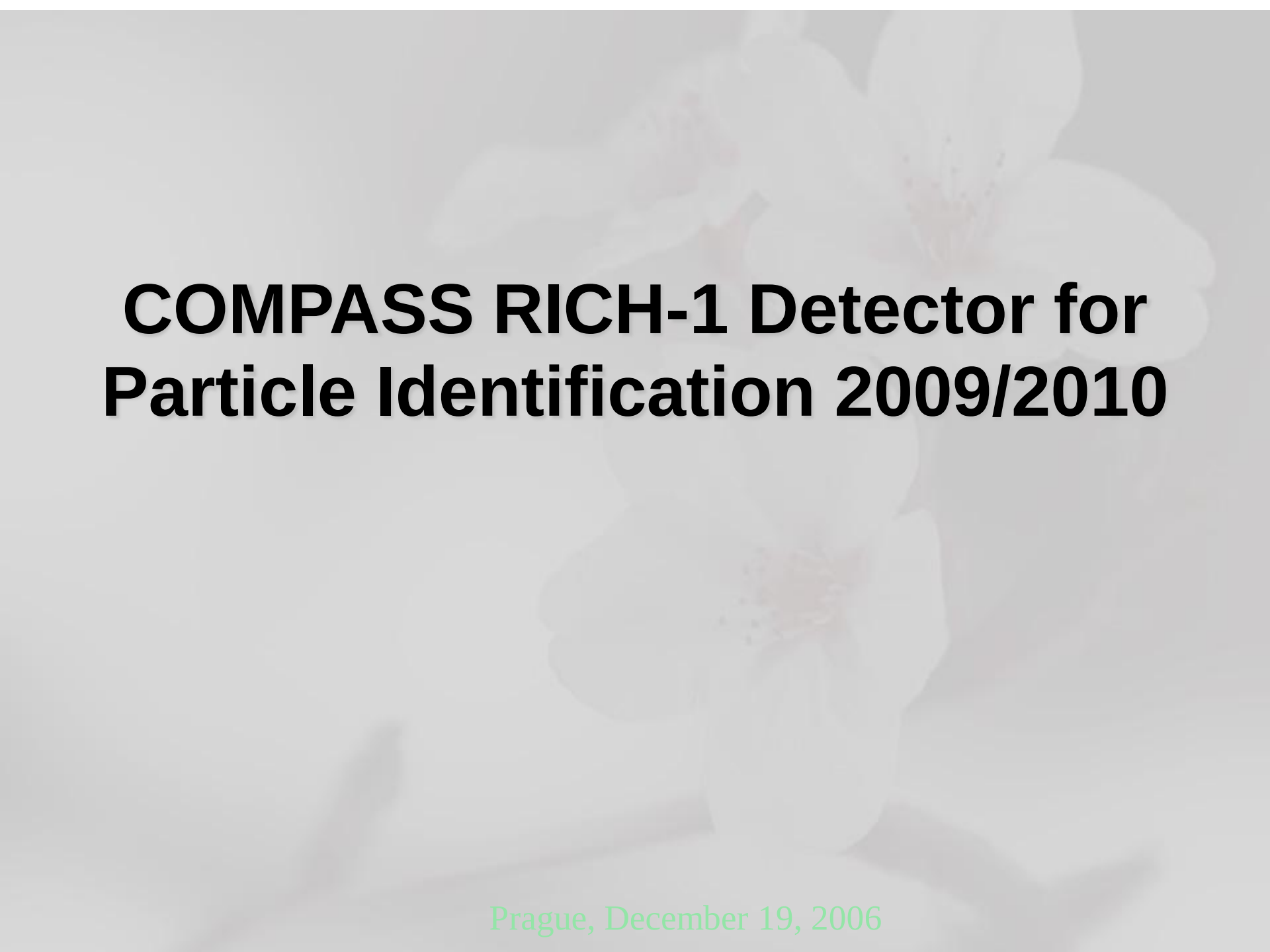
## How to select $n$ ?

- Usually: RICH after magnetic field  
⇒ Minimum Momentum is  $\pi^\pm$  Threshold
- But also reverse:  
CERES RICH
- Use pressure or mixture
- Carefull with sensitive range: Chromatic Dispersion and absorption.

| Material              | $n - 1$              | $\gamma_{\text{thres}}$ |
|-----------------------|----------------------|-------------------------|
| Diamond               | 1.42                 | 1.10                    |
| ZnS (Ag)              | 1.37                 | 1.10                    |
| Lead Fluoride         | 0.80                 | 1.20                    |
| Glass                 | 0.46-0.75            | 1.22-1.5                |
| Water                 | 0.33                 | 1.52                    |
| Aerogel               | 0.025-0.075          | 4.5-20                  |
| CO <sub>2</sub> (STP) | $430 \times 10^{-6}$ | 34.1                    |
| N <sub>2</sub> (STP)  | $300 \times 10^{-6}$ | 45                      |
| Ne (STP)              | $65 \times 10^{-6}$  | 90                      |
| He (STP)              | $33 \times 10^{-6}$  | 123                     |

## Detection of internally reflected Cherenkov light



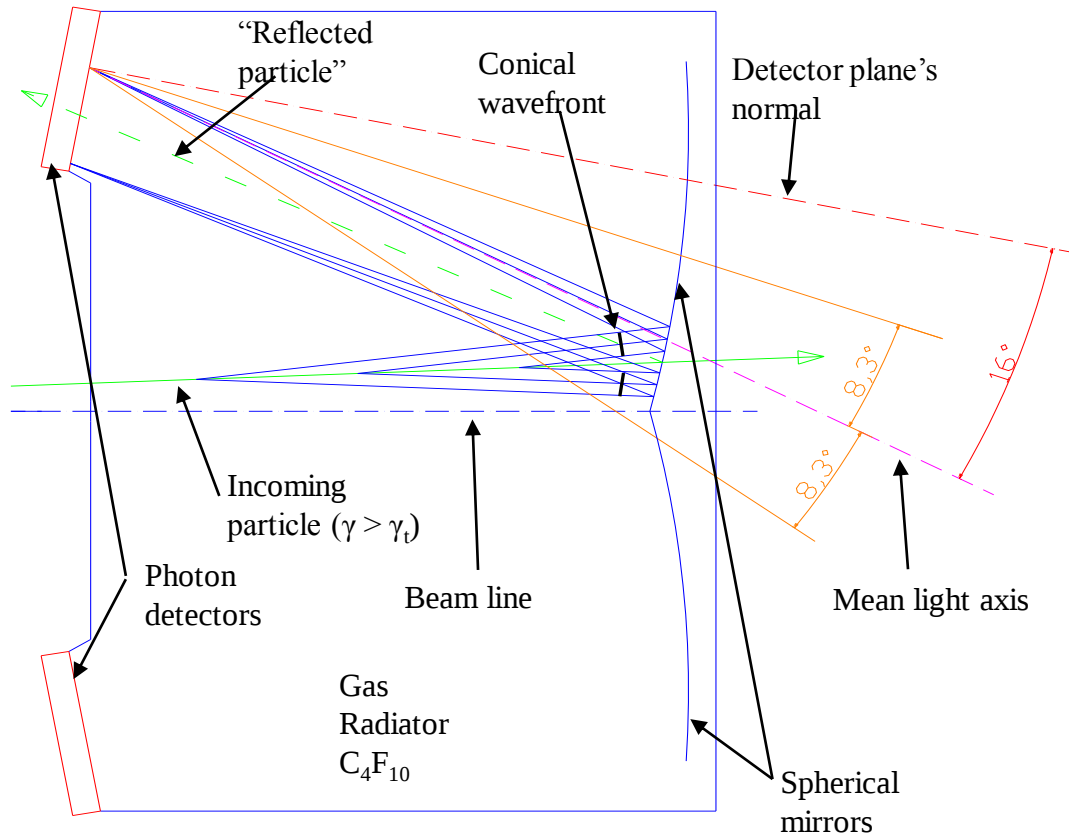
The background of the slide features a faint, artistic image of several white flowers, possibly cherry blossoms, with visible stamens and petals, set against a light gray background.

# **COMPASS RICH-1 Detector for Particle Identification 2009/2010**

Prague, December 19, 2006



# COMPASS *RICH* Detector



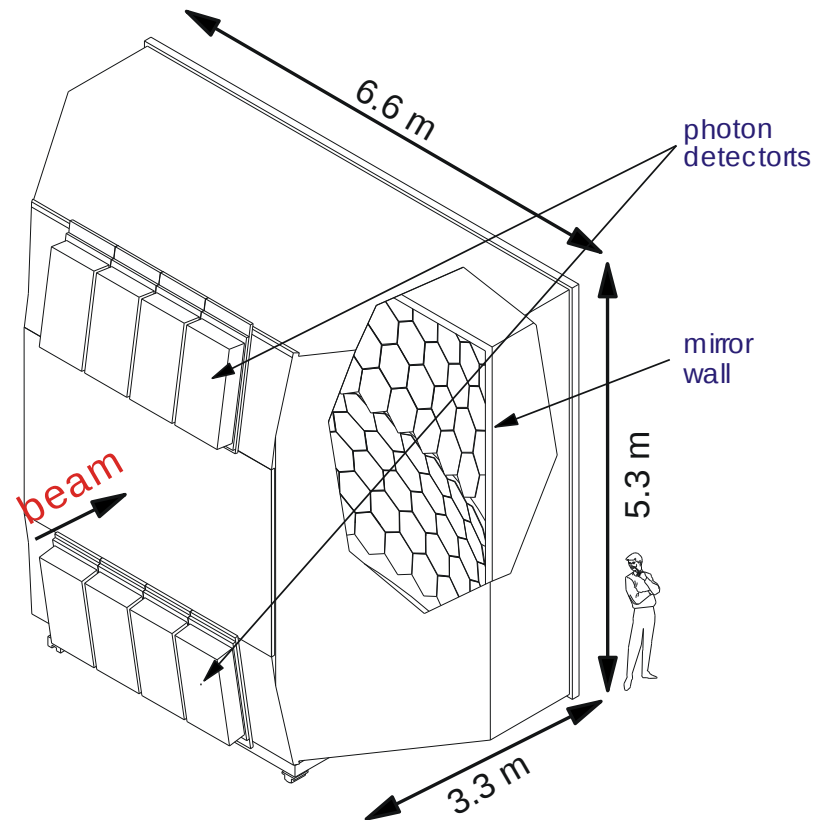
Osaka, APSE2010, June 17, 2010



## RICH-1

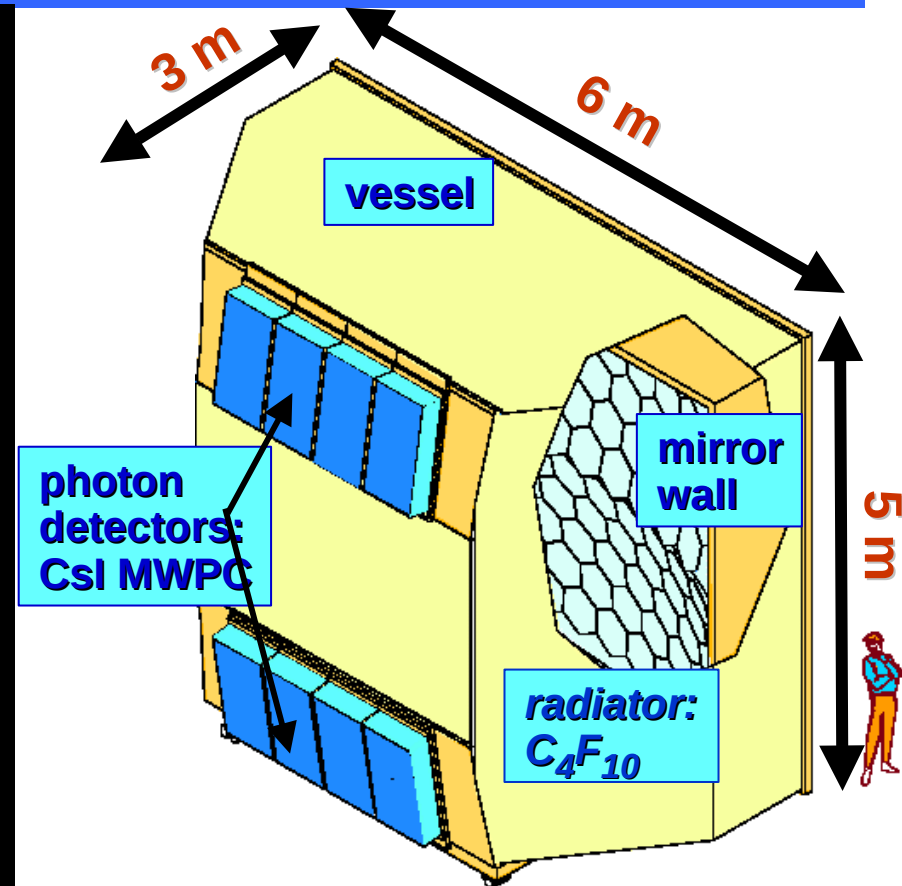
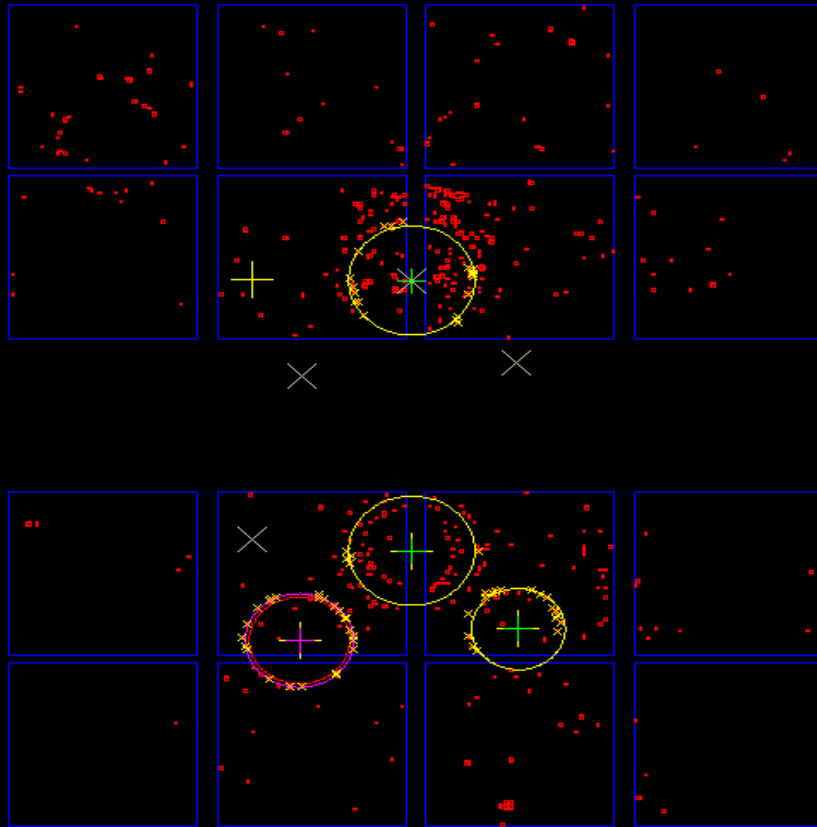
### *Ring Imaging Cherenkov*

- 80 m<sup>3</sup> (3 m C<sub>4</sub>F<sub>10</sub>)
- 116 VUV mirrors (3.3 m focal length)
- 5.3 m<sup>2</sup> VUV detectors
  - MWPC CsI photon-sensitive cathodes
  - 8x8 mm<sup>2</sup> pads
- 84k channels of analog read-out





# COMPASS



Photon detection

5.3 m<sup>2</sup> MWPCs

16 CsI Photocathodes

84,000 analog readout channels

M. Finger

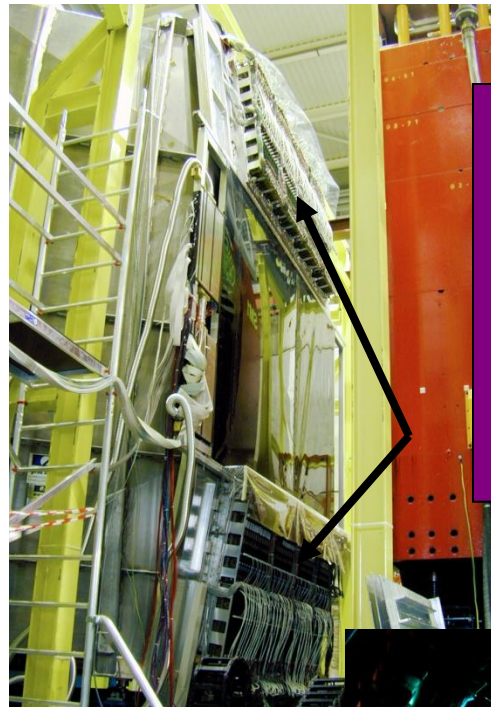
single photon:  $\sigma = 1.2 \text{ mrad}$   
 ring:  $\sigma = 0.4 \text{ mrad}$   
 photons/ring  $n \sim 14$   
 $3\sigma \pi / K \text{ sep.}$  up to 40 GeV/c

Osaka, APSE2010, June 17, 2010



# COMPASS

## RICH-1

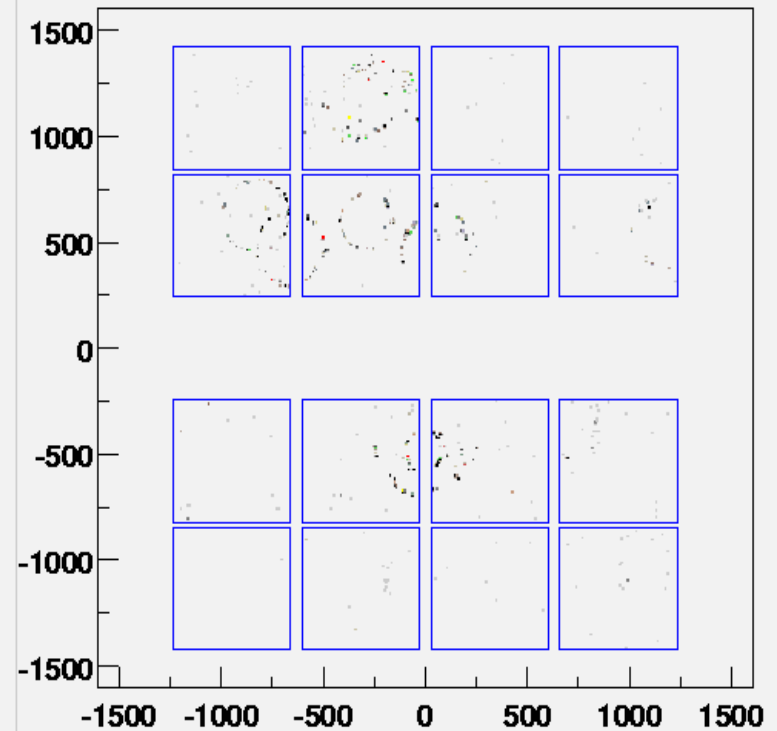


**Photon  
detectors (PD) :**  
MWPCs with CsI  
photocathodes  
(5.3 m<sup>2</sup>),  
84,000 analog  
read-out ch.s

**VUV mirror  
wall, 21 m<sup>2</sup>,  
116 mirrors**



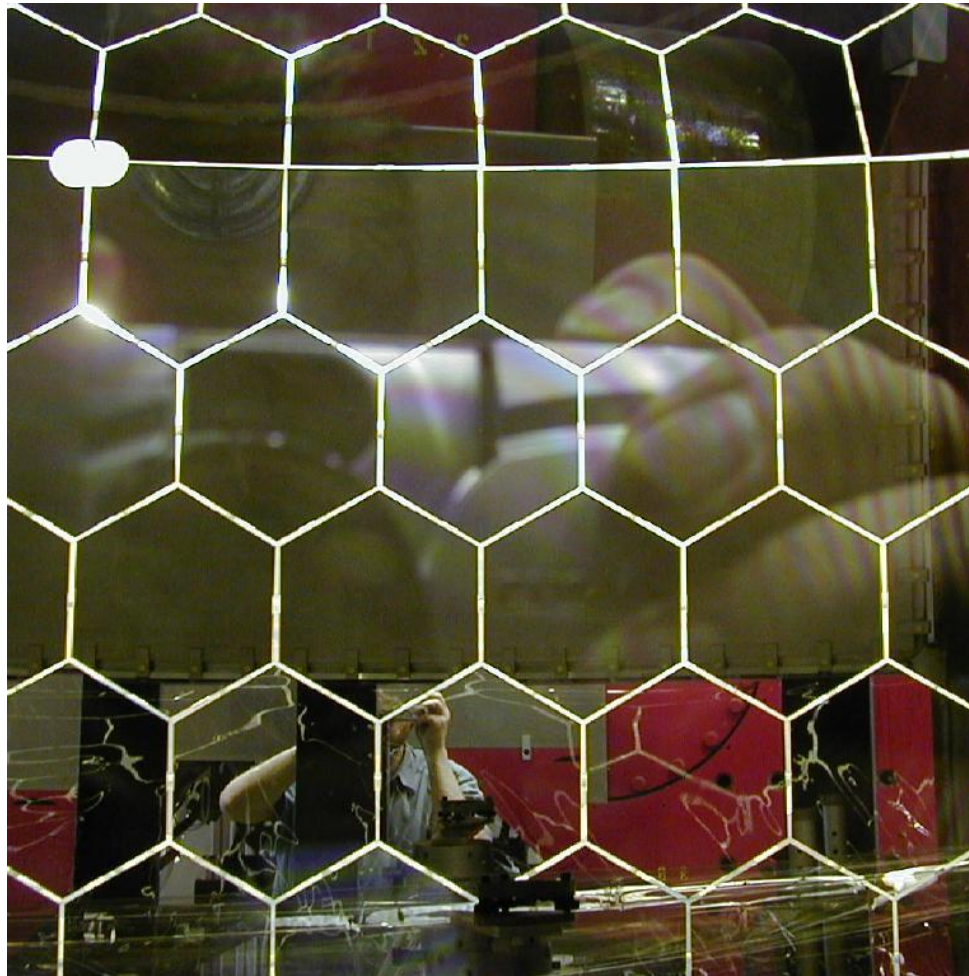
**an event from  
on-line display**



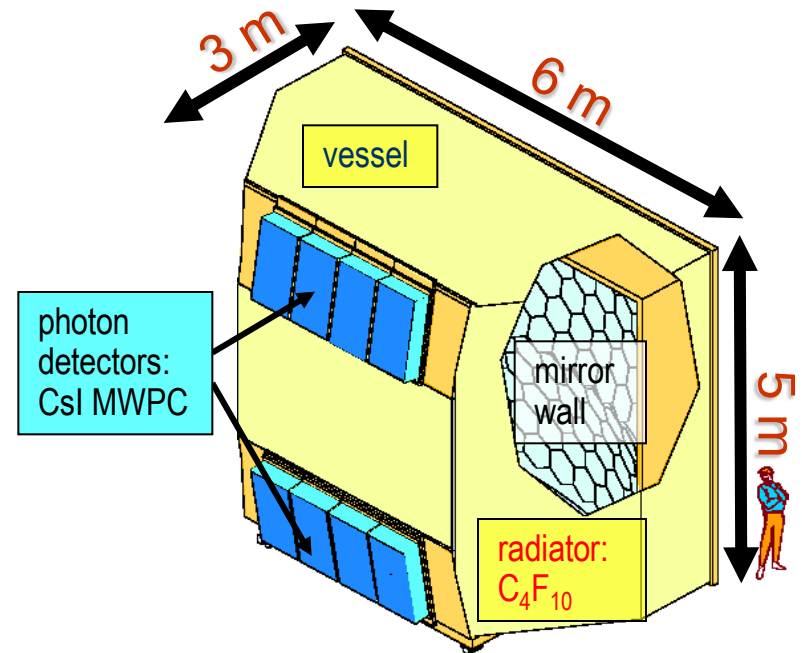
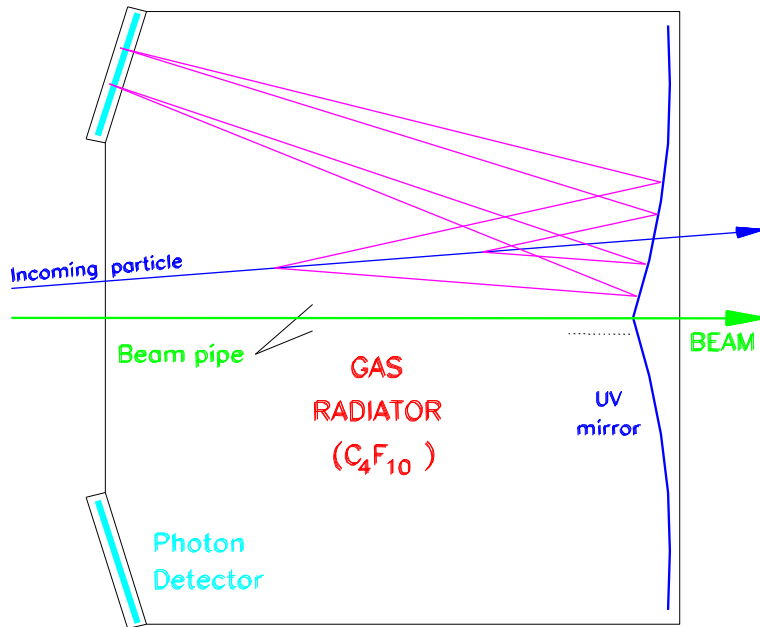


# COMPASS

## RICH-1 Mirror Wall



# The COMPASS RICH detector (2003-2004, before upgrade)



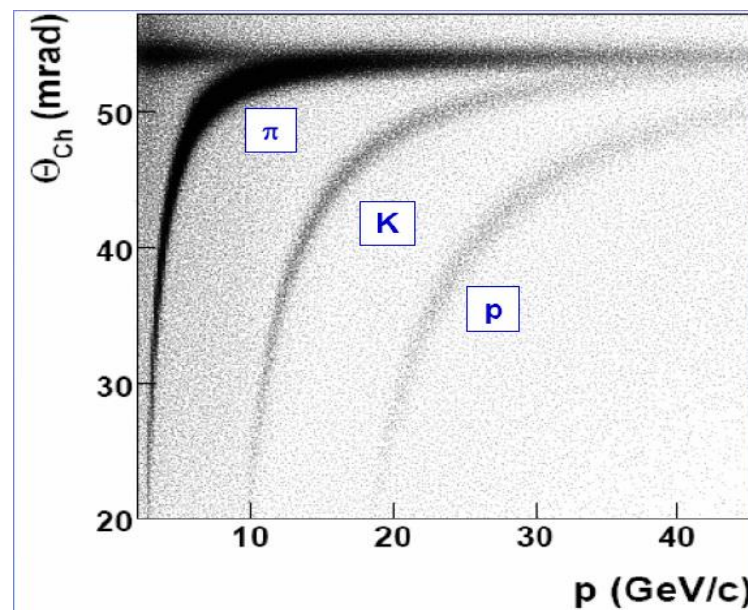
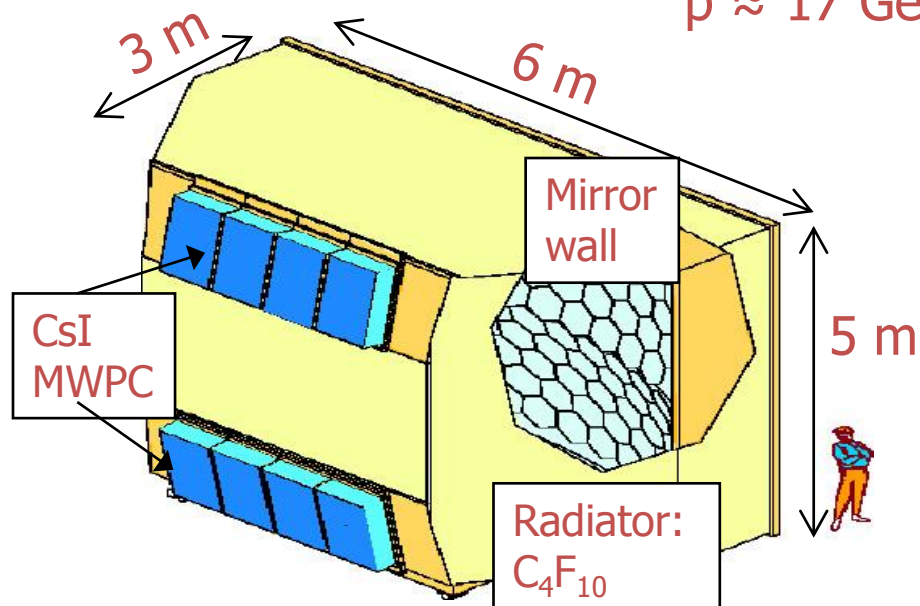
- **radiator gas:** C<sub>4</sub> F<sub>10</sub>, good transparency > 160 nm, low chromaticity
- **mirror:** 20 m<sup>2</sup> surface, VUV reflectance (160-210 nm) > 80%
- **photon-detectors:** Multi Wire Proportional Chamber, total surface 5.3 m<sup>2</sup>
- **angular acceptance:** ± 250 mrad horizontal, ± 200 mrad vertical
- **read-out:** 83.000 channels (pixels)

# particle identification with Compass RICH

identification of  $\pi$ , K and protons

Cherenkov thresholds:  $\pi \approx 3 \text{ GeV/c}$   
 $K \approx 9 \text{ GeV/c}$   
 $p \approx 17 \text{ GeV/c}$

$2\sigma$   $\pi/K$  separation at  $43 \text{ GeV/c}$



|       | without<br>PID   | $\pi^+ \pi^-$    | $\pi^+ K^-$      | $K^+ \pi^-$      | $K^+ K^-$        |
|-------|------------------|------------------|------------------|------------------|------------------|
| total | $5.3 \cdot 10^6$ | $3.7 \cdot 10^6$ | $2.4 \cdot 10^5$ | $3.0 \cdot 10^5$ | $8.7 \cdot 10^4$ |

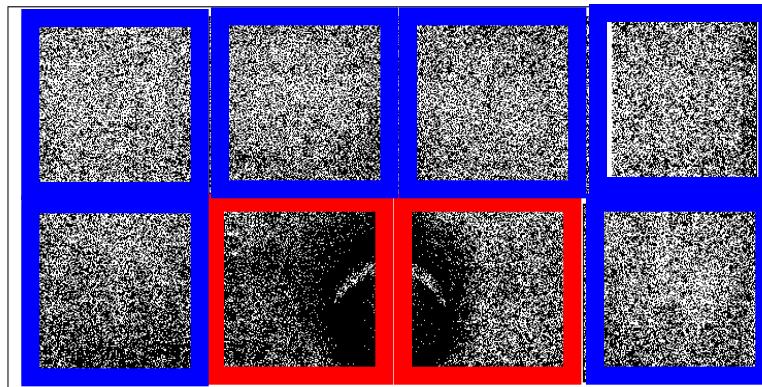
# The upgrade project for 2006 data taking

## Multi Wire Proportional Chambers:

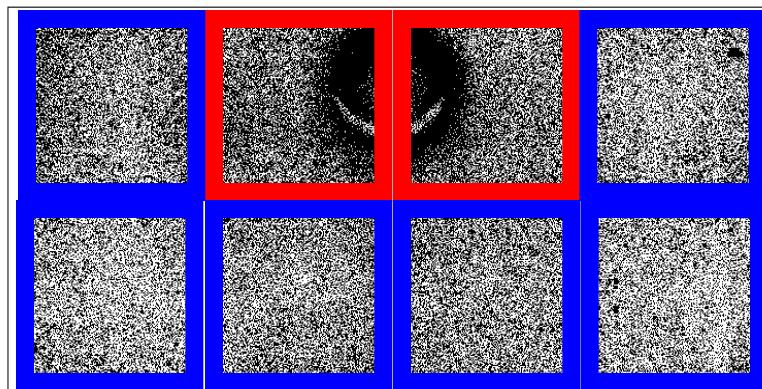
(MWPC, in use since 2001)

- CsI photo cathodes
- new read-out system: APV chip (negligible dead time)
- time resolution (CsI with MWPC + ro):  
 $\sim 3 \mu\text{s} \rightarrow < 40 \text{ ns}$

Upper Chambers



Lower Chambers



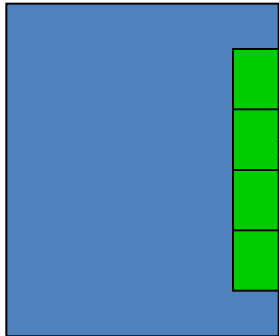
## Fast photon detection:

- Multi Anode PMTs (MAPMT, 576 pc)
- individual telescopes
- read-out
  - sensitive FE: MAD chip
  - high resolution TDC: F1 chip
- time resolution: few ns
- no MWPC

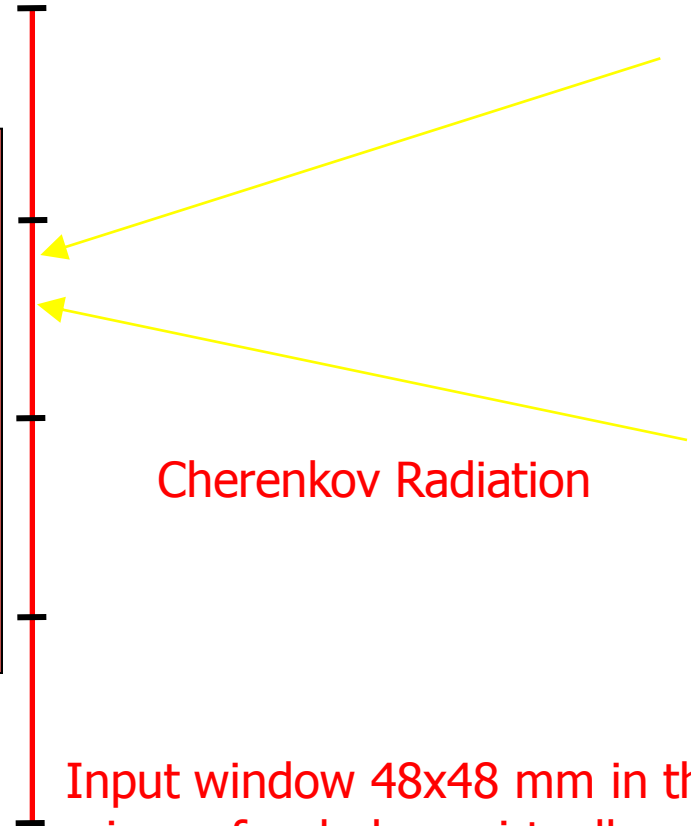
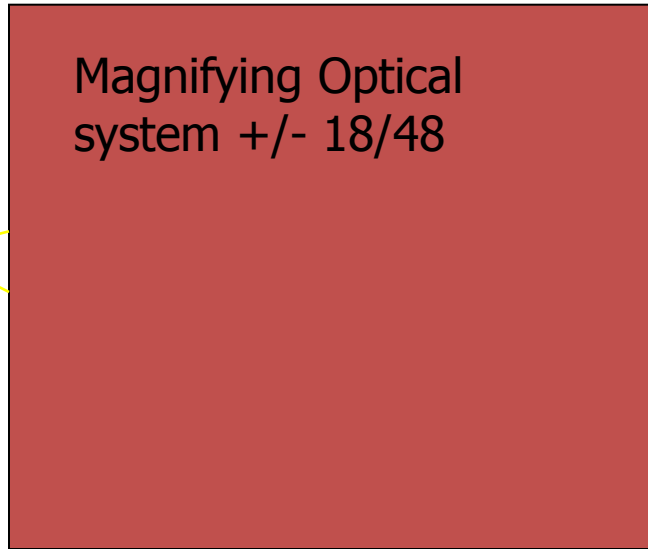
# Optical system for upgrade RICH-1 detector

4x4 Multichannel  
Photomultiplier

with active region  
18x18 mm



Magnifying Optical  
system +/- 18/48



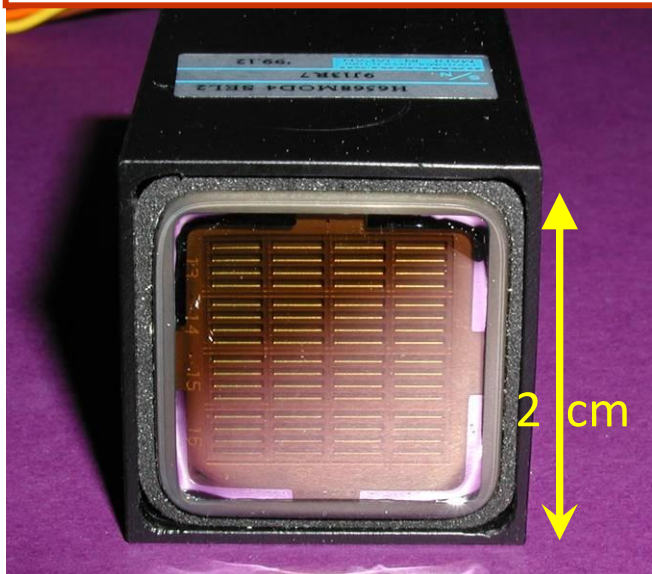
Cherenkov Radiation

Input window 48x48 mm in the  
mirrors focal plane virtually  
divided into 4x4 regions

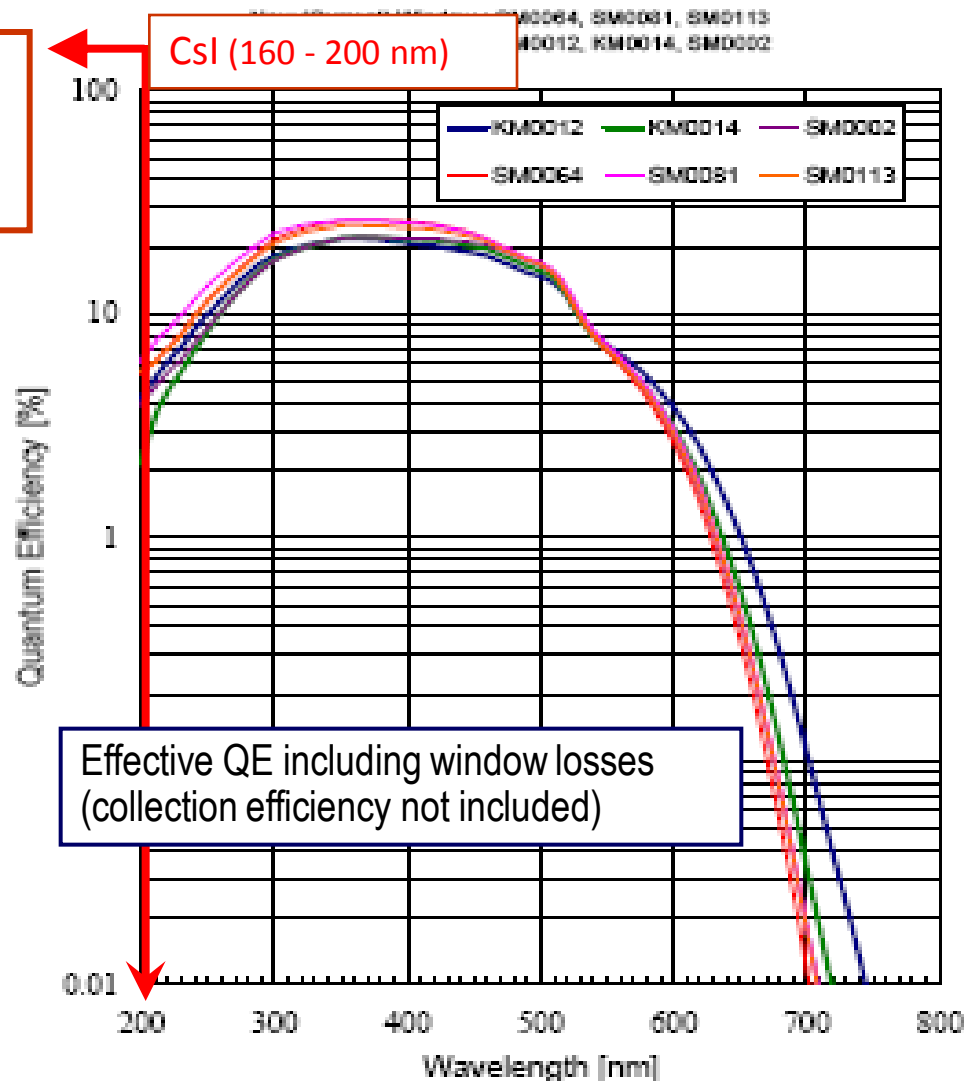
# Hamamatsu R7600-03-M16 photomultiplier

## R7600-03-M16 Spectral Response Characteristics

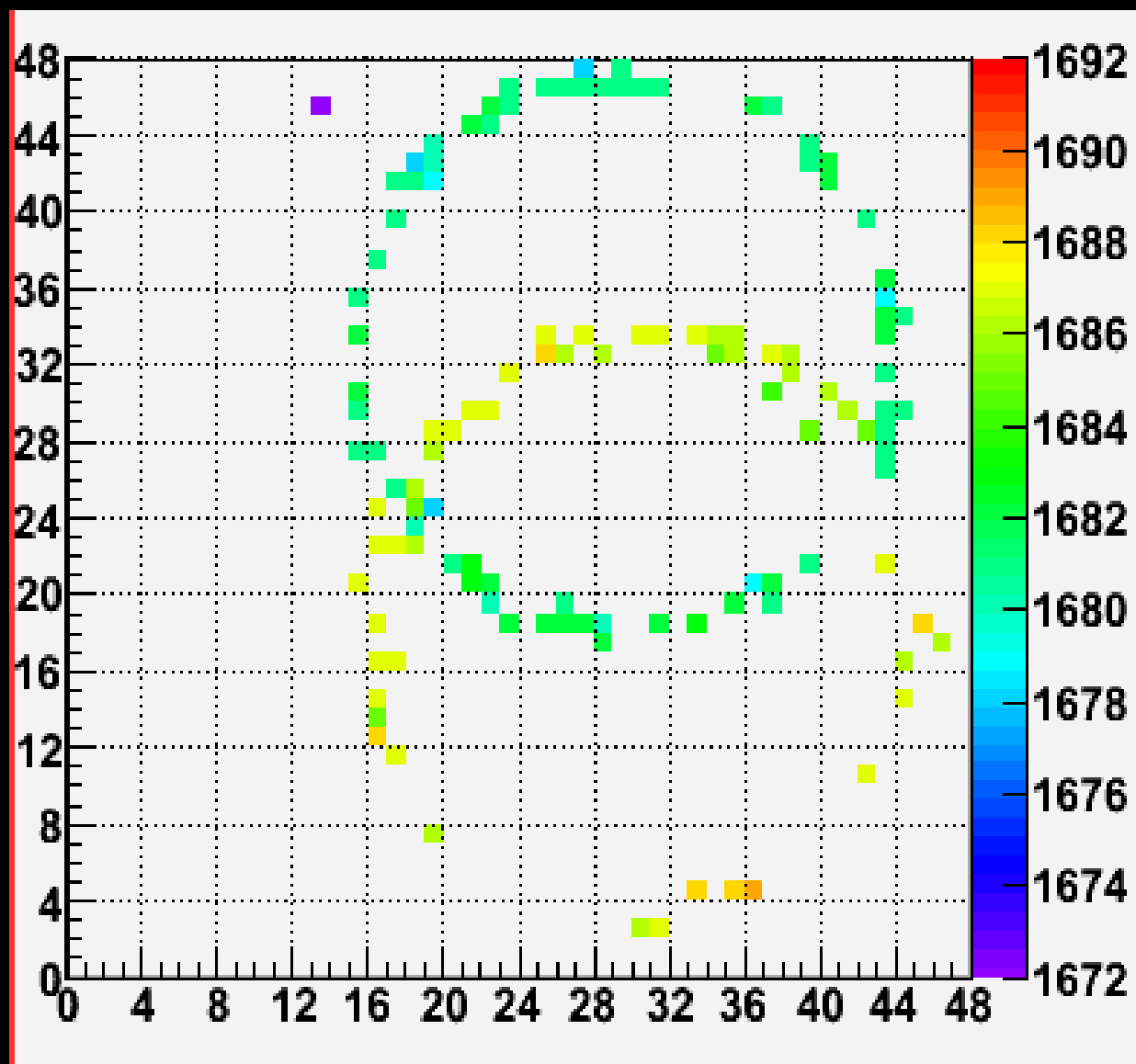
Larger number of detected Cherenkov photons, due to **larger wavelength range**:



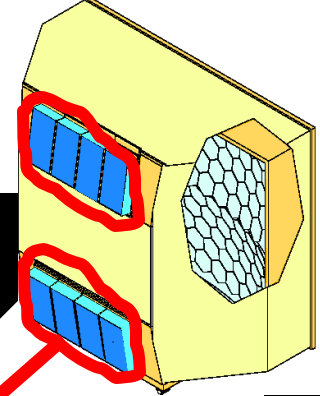
- alkali photocathode
- 18 x 18 mm<sup>2</sup> active area
- 16 pixel
- time resolution 300 ps
- UV transparent borosilicate glass window



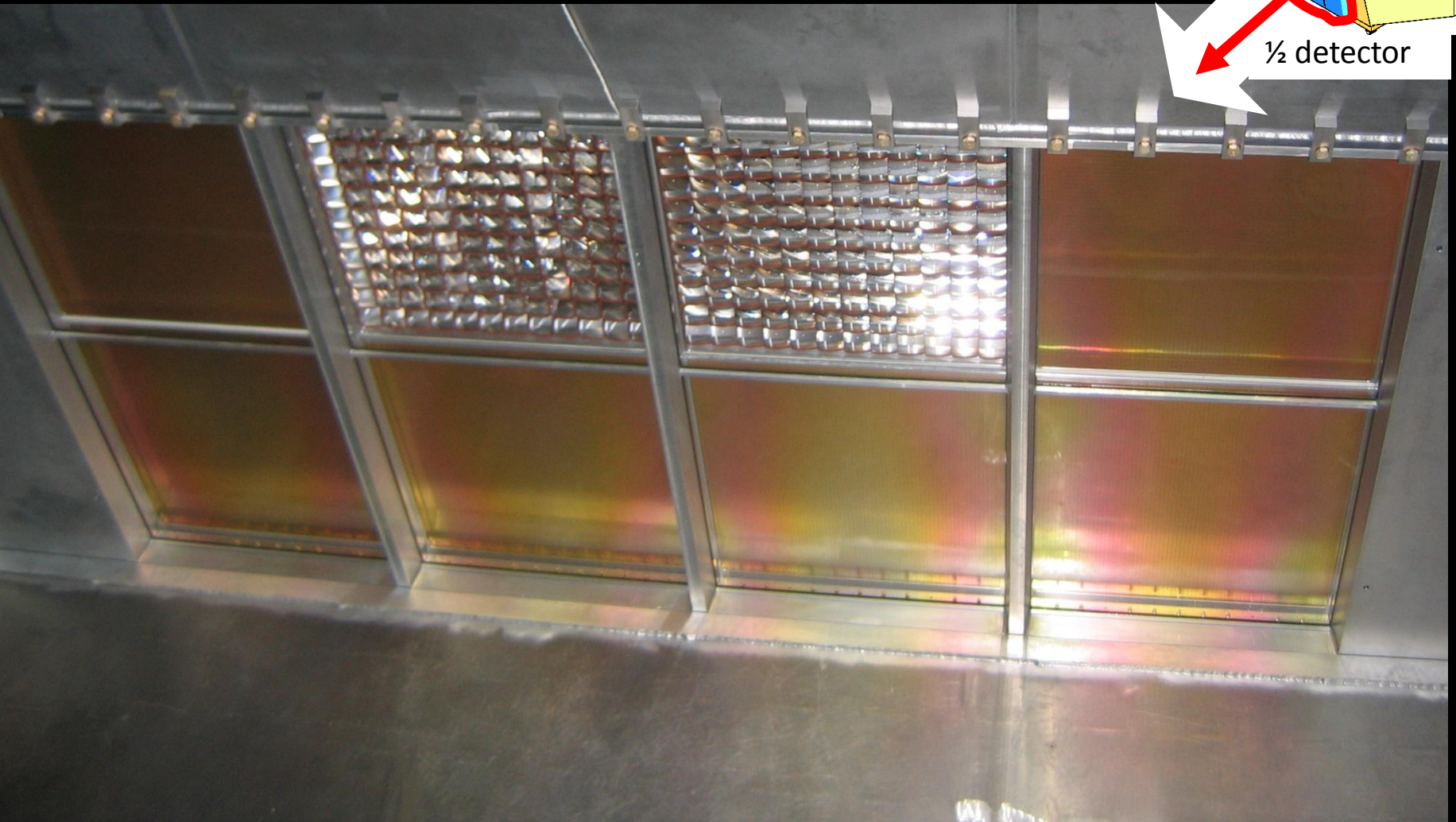
Time resolution is useful for correctly assigning hits to rings:



## Half upgraded detector- from inside



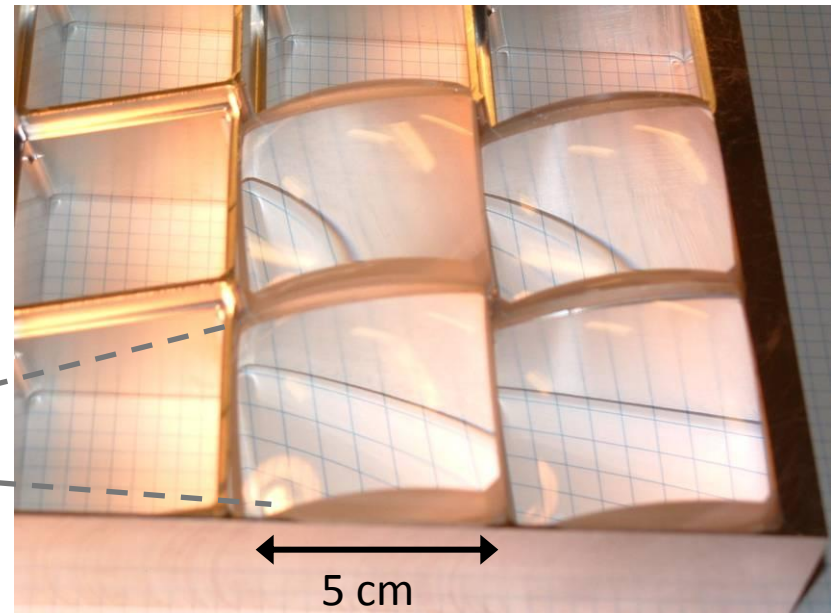
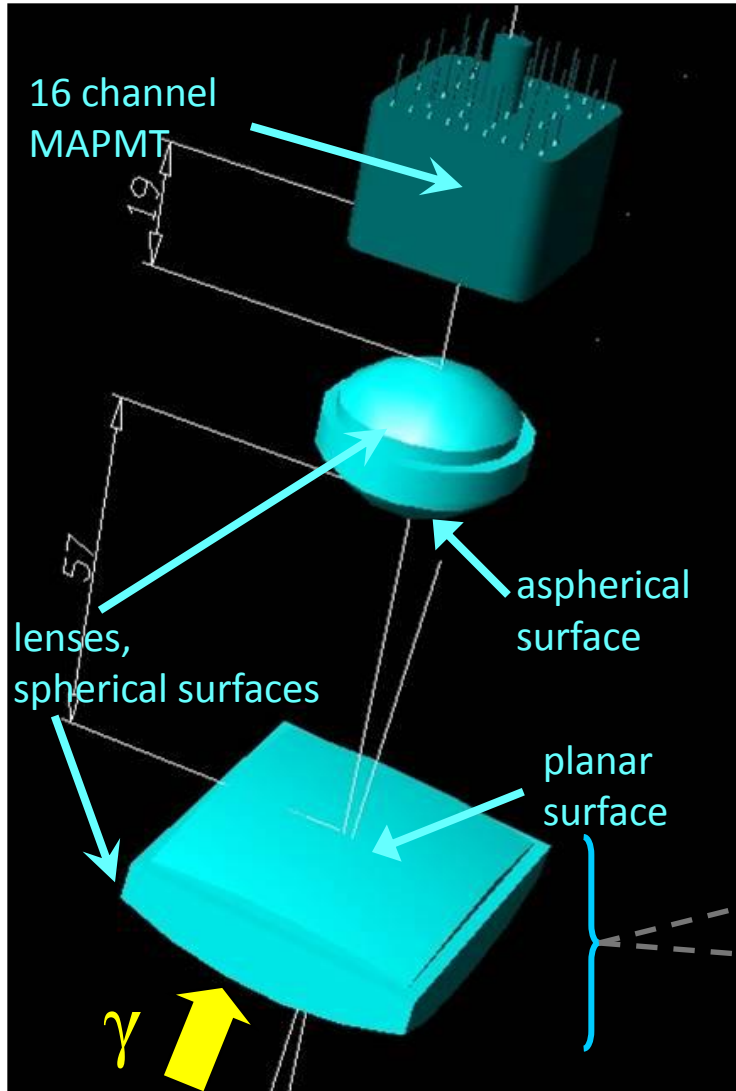
$\frac{1}{2}$  detector



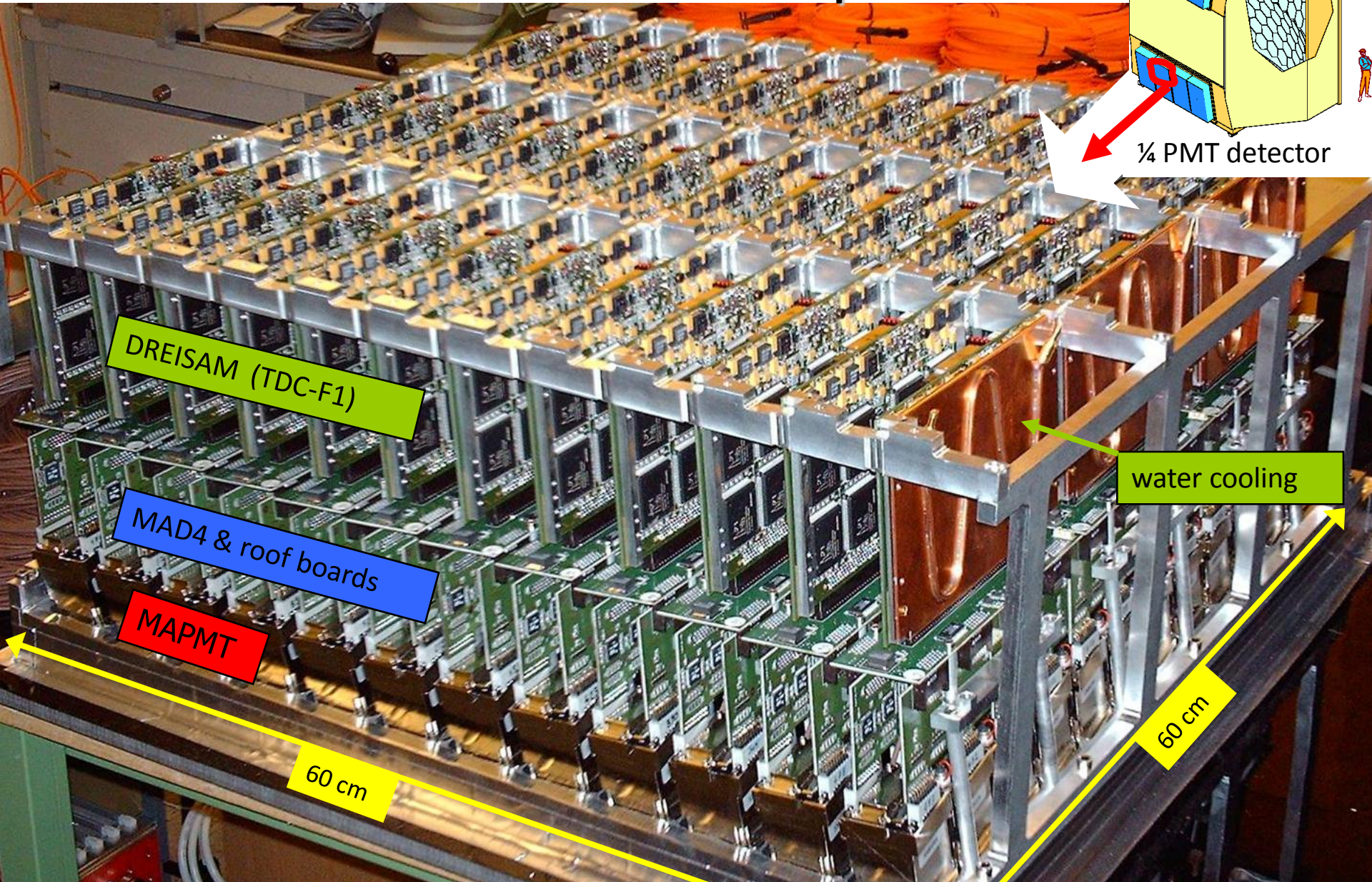
# 576 telescopes made of silica lenses

## Telescope system consisting of 2 lenses:

- Purpose: Focussing Cherenkov photons on MAPMTs (factor 7)
- UV transparent quartz lenses
- large geometrical acceptance
- minimum image distortion
- optimised by Zemax simulation



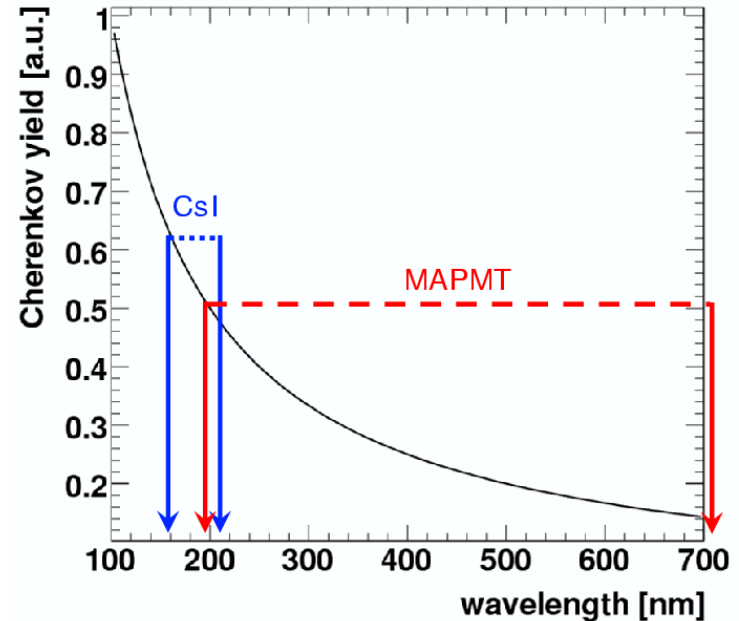
# Read-out electronics of 1st quarter



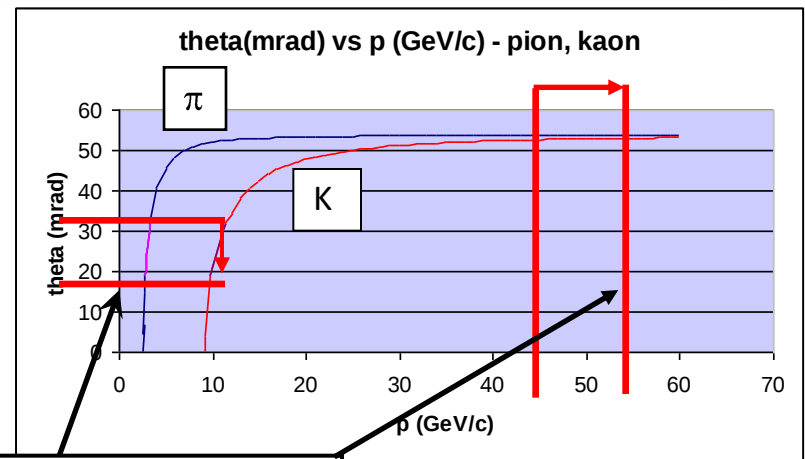
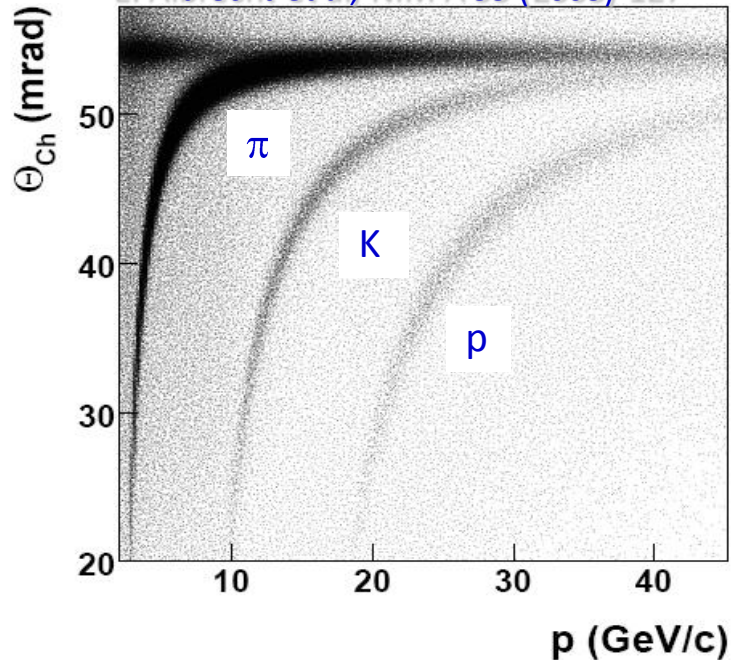
# Upgraded RICH resolution

## Expected performances:

- $\sigma_{ph} = 2.4 \text{ mrad}$  ( $\beta \approx 1$ ) - before: 1.2 mrad
- $N_{ph}/ring \approx 50$  ( $\beta \approx 1$ ) - before: 14
- $\sigma_{ring} \approx 0.4 \text{ mrad}$  ( $\beta \approx 1$ )  
- before: 0.6 mrad
- $2\sigma$   $\pi/K$  separation at  $p \approx 50 \text{ GeV/c}$   
- before: 43 GeV/c



E. Albrecht et al, NIM A 33 (2003) 127



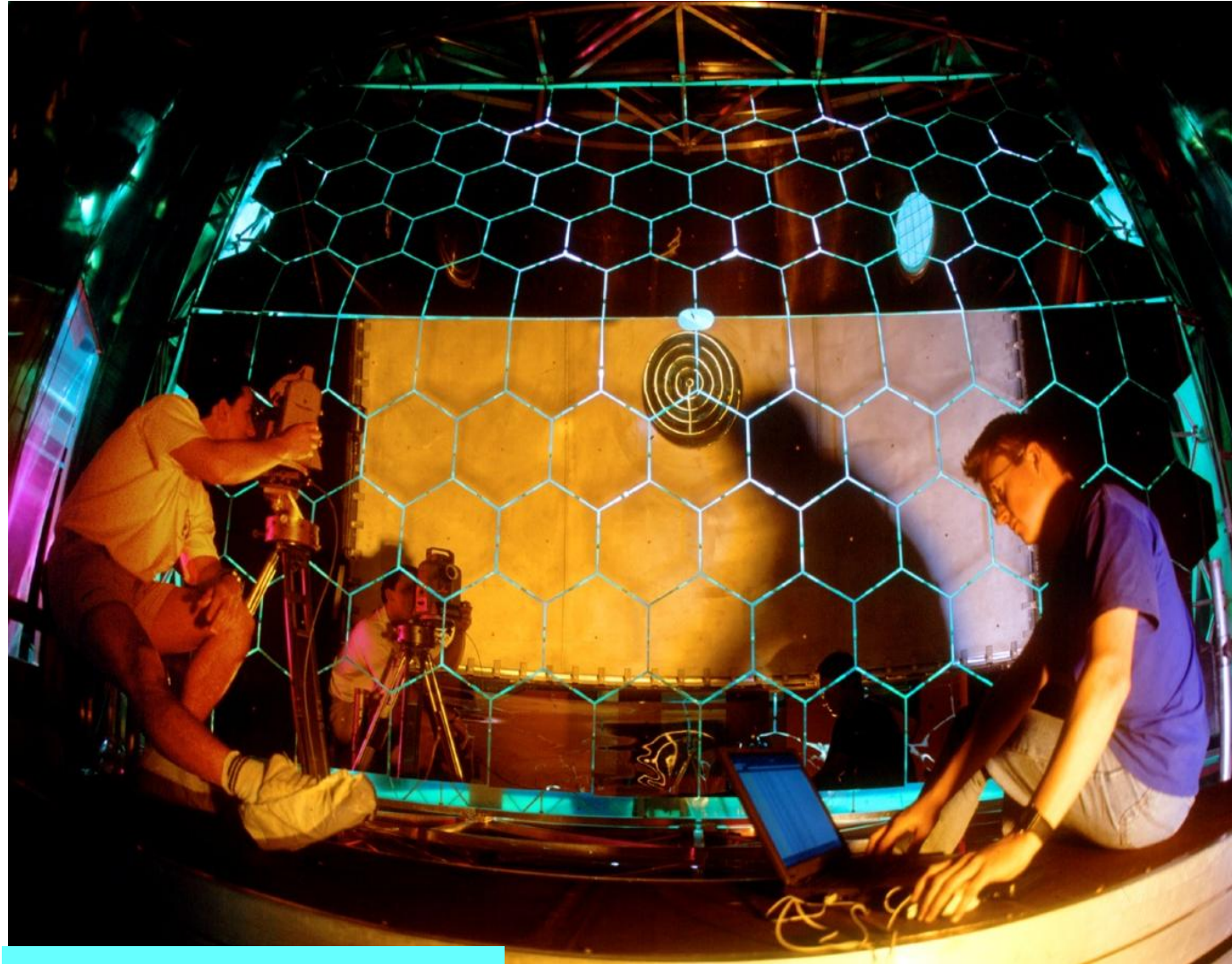
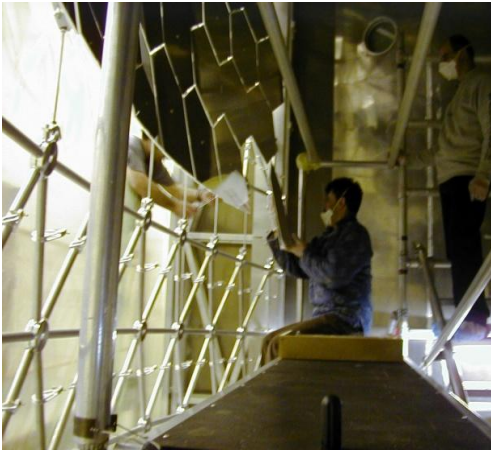
PID capability  
extended by upgrade

# Performances of upgraded RICH detector

Excellent performances:

- $N_{\text{ph}}/\text{ring}$  ( $\beta \approx 1$ )  $> 60$  before upgrade: 14
- time resolution  $< 1 \text{ ns}$  before upgrade:  $3 \mu\text{s}$
- $\sigma_{\text{ring}} \approx 0.3 \text{ mrad}$  ( $\beta \approx 1$ ) before upgrade:  $0.6 \text{ mrad}$
- $N_{\text{ph}}/\text{ring} \approx 50$  ( $\beta \approx 1$ ) - before: 14
- $2\sigma$   $\pi/K$  separation at  $p_h > 55 \text{ GeV}/c$  before upgrade:  $43 \text{ GeV}/c$
- excellent suppression of background from  $\mu$  halo,
- performances as expected (and better)

# mirrors and alignment



21 m<sup>2</sup>, 116 mirrors  
radius: 6.6 m

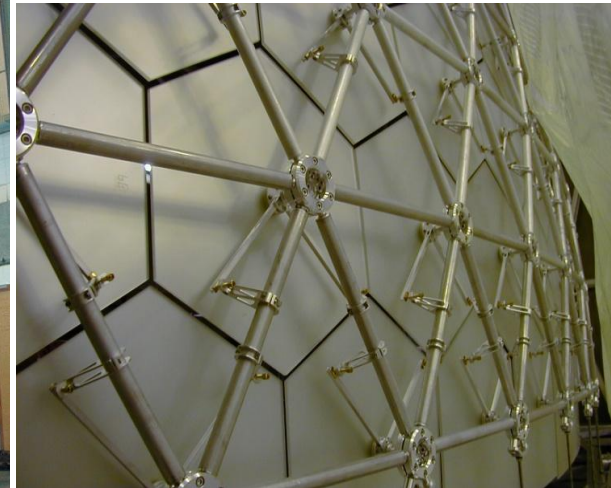
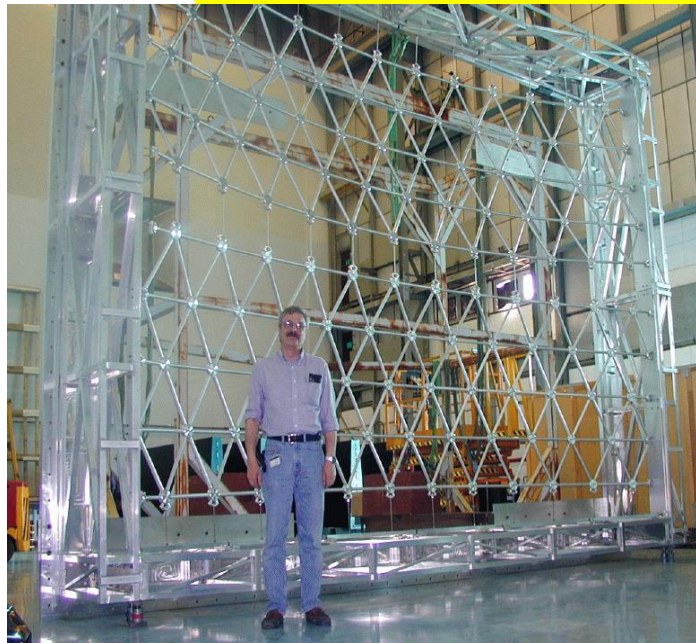
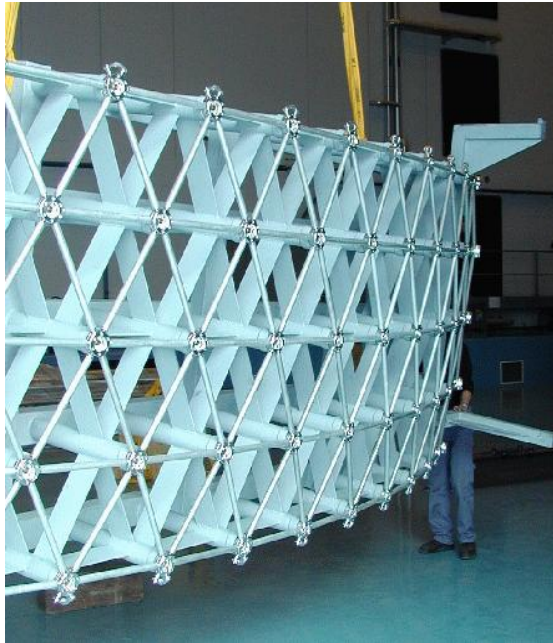
angular regulation screws

measurement of mirror alignment  
via laser autocollimation

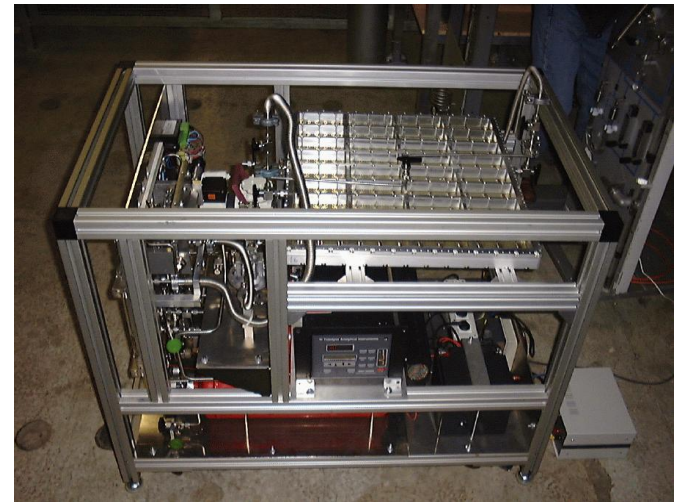
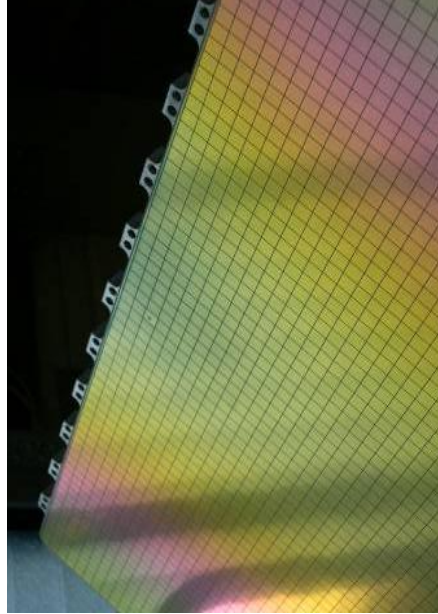
# the vessel and the mirror support



Large and accurate mechanics  
light front and rear windows



# The CsI photocathodes



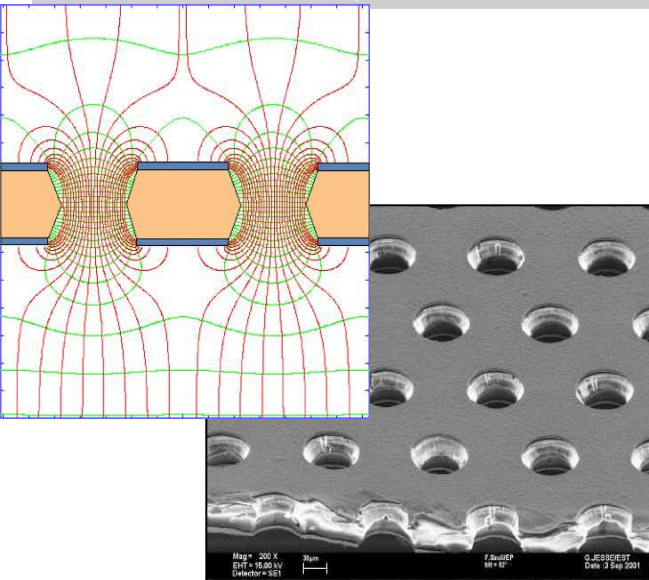
# There is need of new technology

*to overcome recent limits – fight ion bombardment and photon feedback*

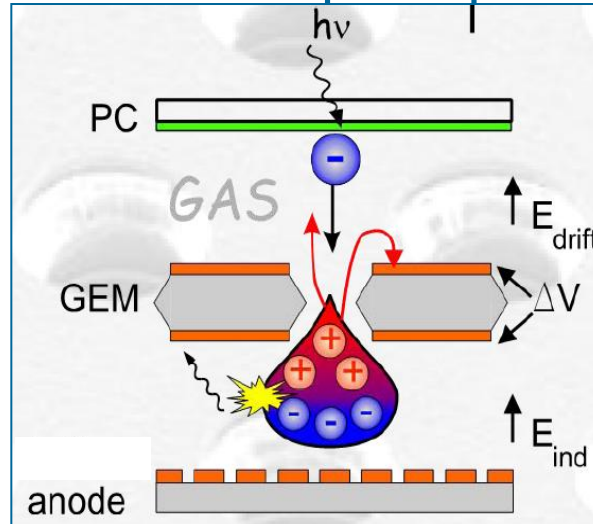
*Possible solution – closed geometries*

## GEMs and THGEMs

GAS ELECTRON MULTIPLIER (GEM)  
Thin metal-coated polymer foils  
70  $\mu\text{m}$  holes at 140  $\mu\text{m}$  pitch

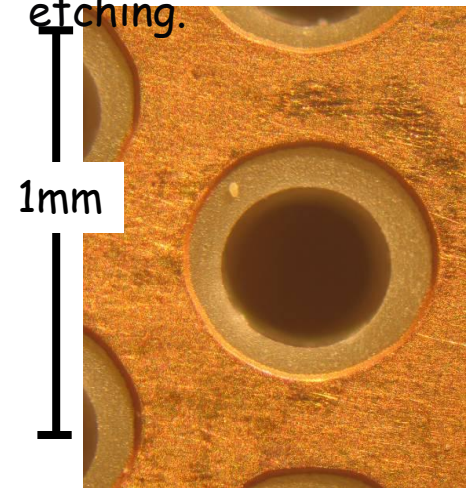


### GEM's principle



### ECONOMIC & ROBUST

Manufactured by  
**standard PCB techniques**  
of precise drilling and Cu  
etching.



F. Sauli, NIM A386(1997)531

Chechik et al. NIM A535 (2004) 303  
C. Shalem et al. NIM A558 (2006) 475  
& NIM A558 (2006) 468

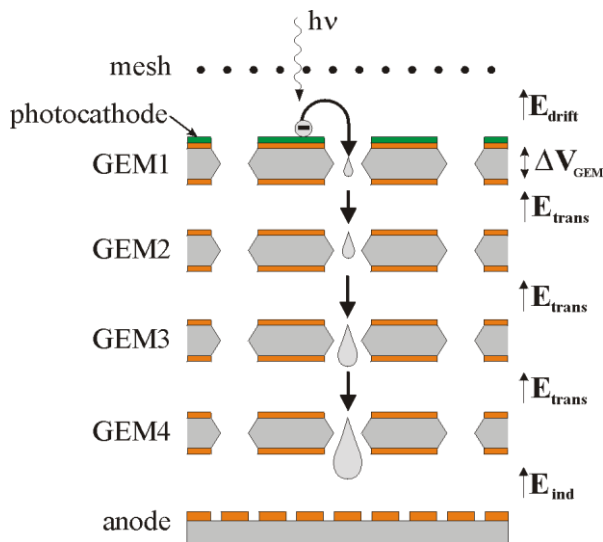
# MultiGEM's: ion blocking + high GAIN

Examples of ion blocking schemes from literature

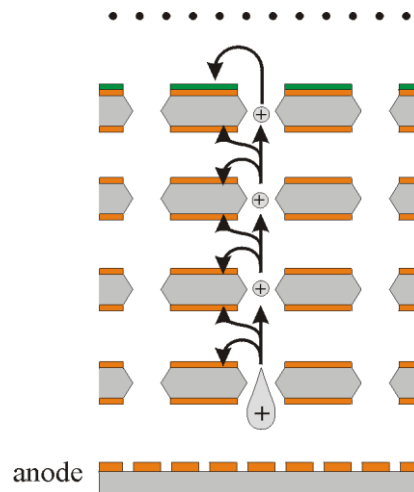
- Similar schemes can be adopted with THGEM

## multiGEM with Reflective photocathode

### Electron's path

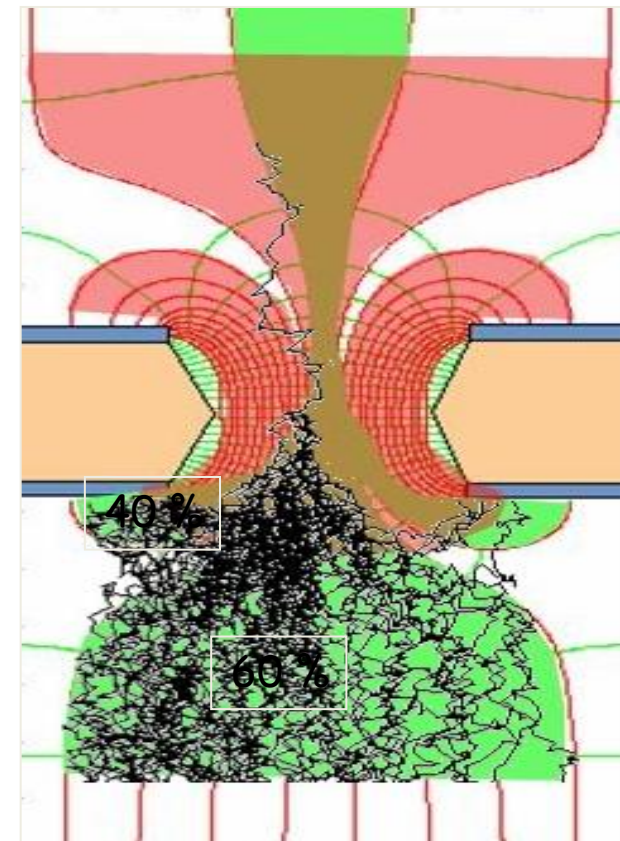


### Ion's back-flow



## Simulation of avalanche

Ions



Electrons

# Perspectives

## Short term plans:

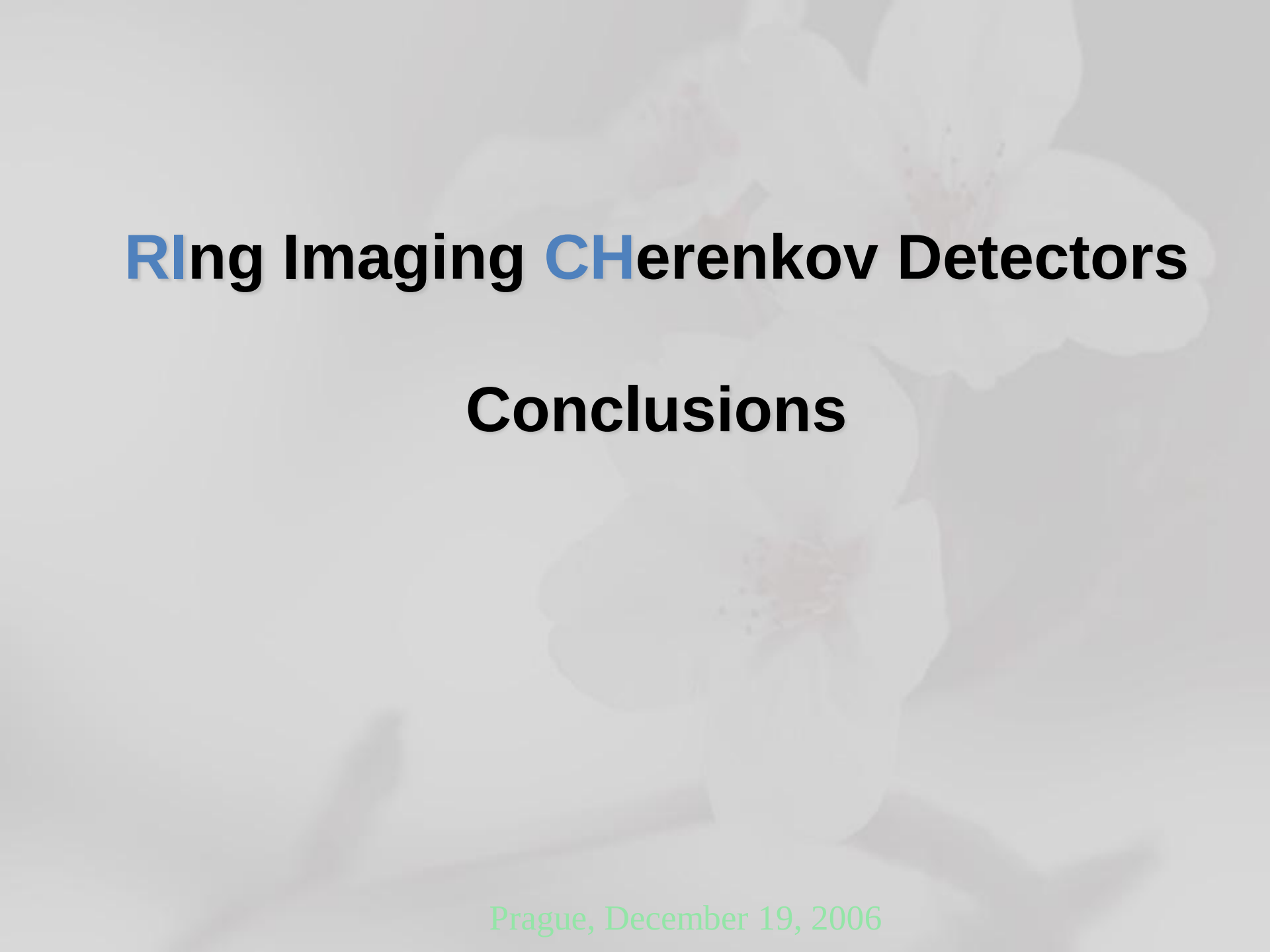
- optimize the parameters of the THGEM with photoconverting CsI layer to achieve maximum photoelectron collection efficiency
- optimize the parameters for the (double) THGEM to be used for the amplification of the signal to provide large and stable gain
- produce a set 600 x 600 mm<sup>2</sup> THGEMs and assemble them with stesalite spacer frames into first complete “full size” prototype chamber

## Possible medium term project:

- Upgrade of COMPASS RICH ( $\sim 4\text{m}^2$ ) with the new photon detectors in case the COMPASS Collaboration decides for it.

## Longer term dream:

- find a configuration to reduce the ion back-flow down to  $<10^{-5}$  and operate this large area detectors with visible photoconverter

The background of the slide features a faint, artistic image of white flowers, possibly cherry blossoms, with visible petals and stamens. The image is light and serves as a subtle backdrop for the text.

# **RI**ng Imaging **CH**erenkov Detectors

## **Conclusions**

Prague, December 19, 2006

# COMPASS RICH-1

is a large gaseous RICH

with two kind of photon detectors

providing:

hadron PID from 3 to 60 GeV/c

acceptance: H: 500 mrad V: 400 mrad

trigger rates: up to  $\sim 100$  KHz

beam rates up to  $\sim 10^8$  Hz

material in the beam region:  $2.4\% X_0$

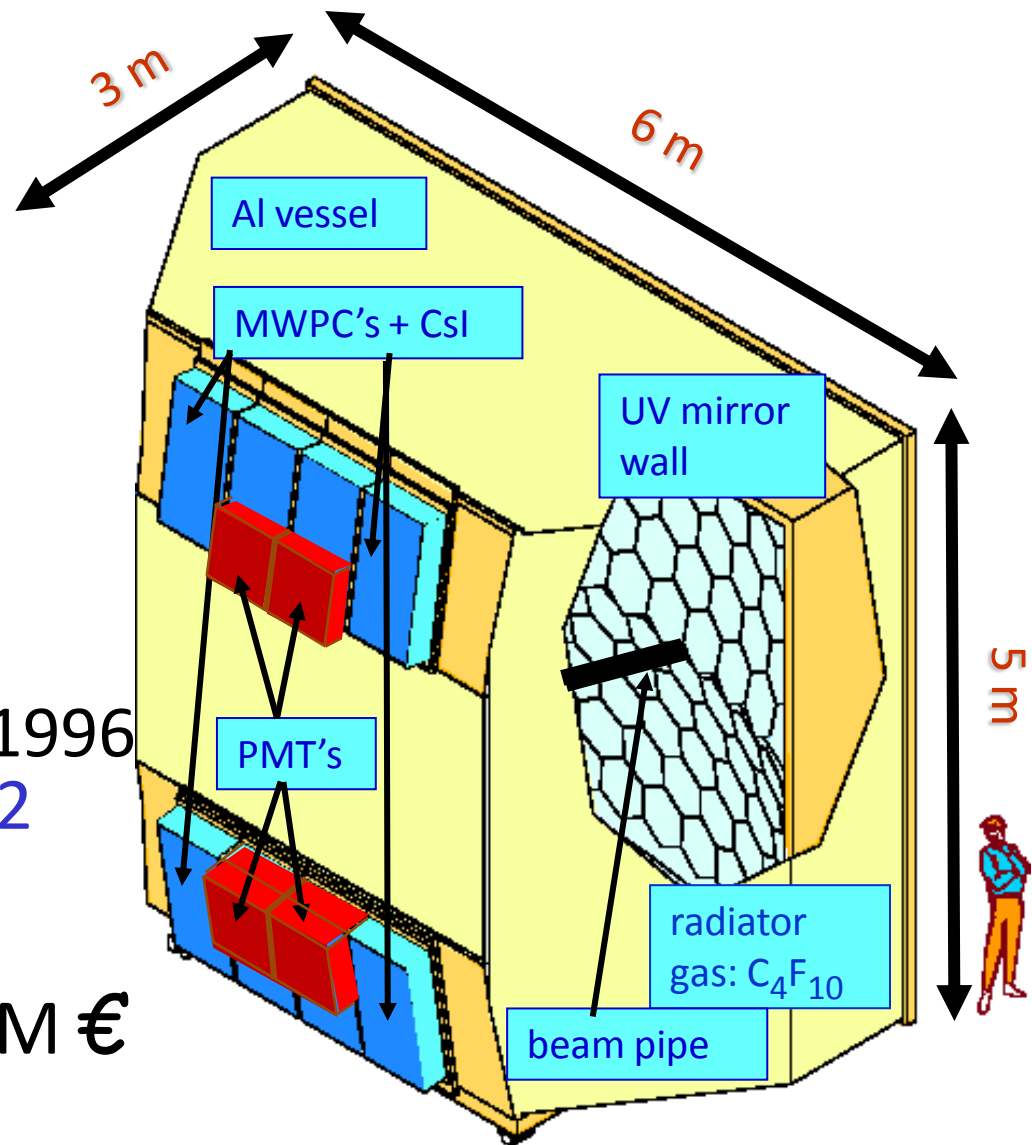
material in the acceptance:  $22\% X_0$

\* detector designed in 1996

\* in operation since 2002

\* upgraded in 2006

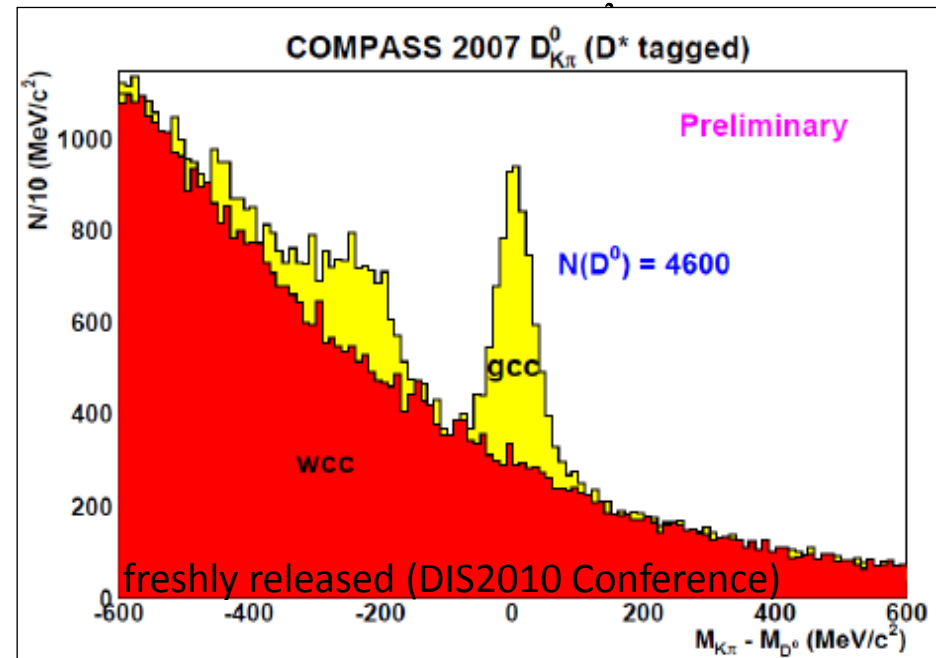
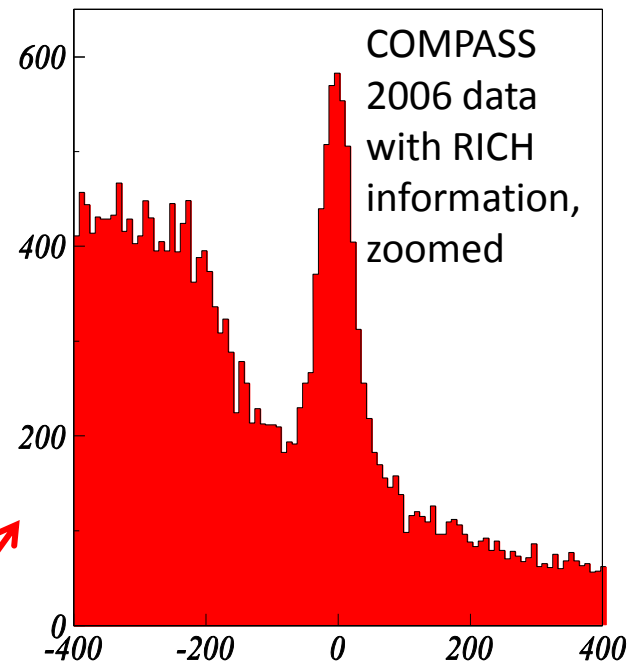
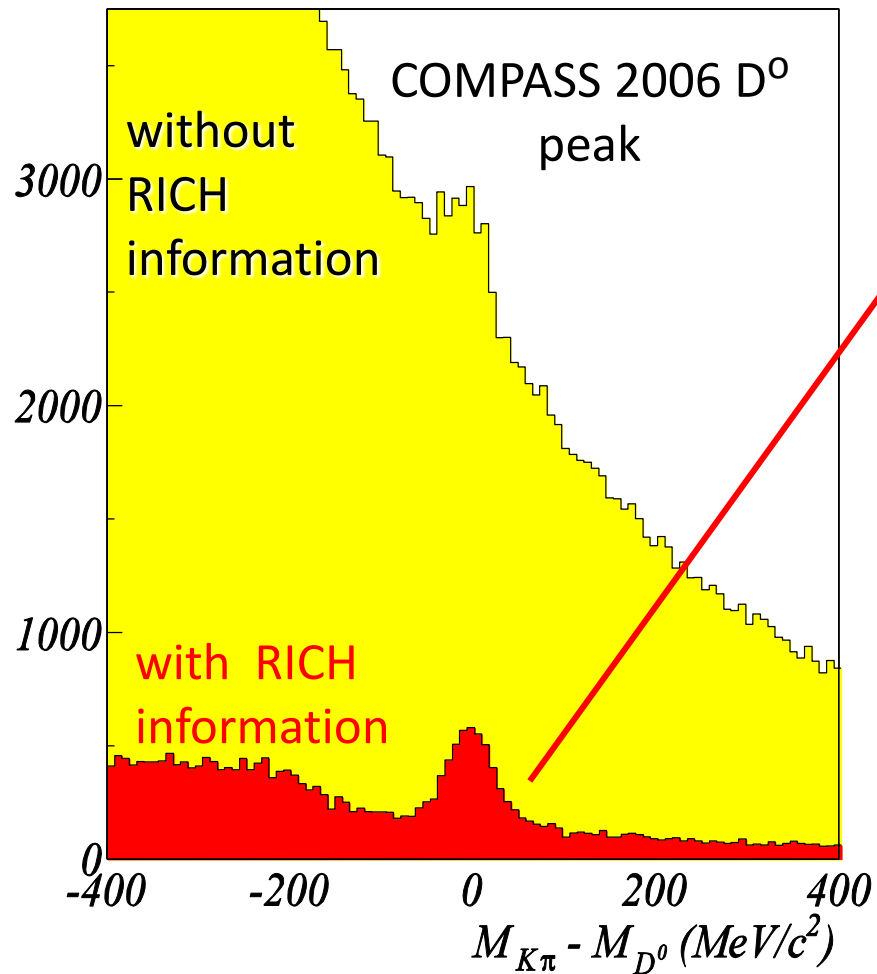
total investment:  $\sim 4$  M €



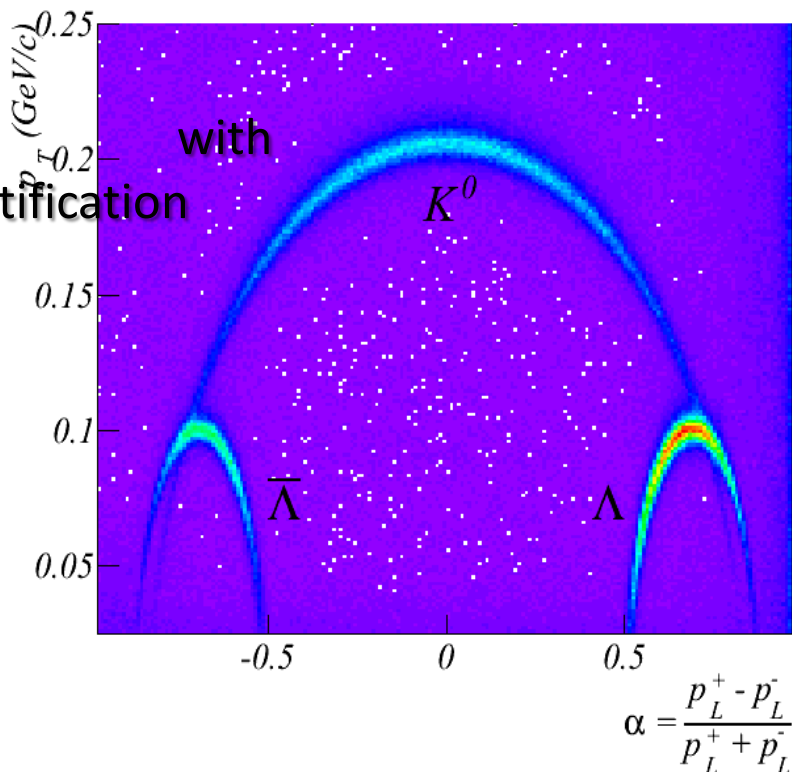
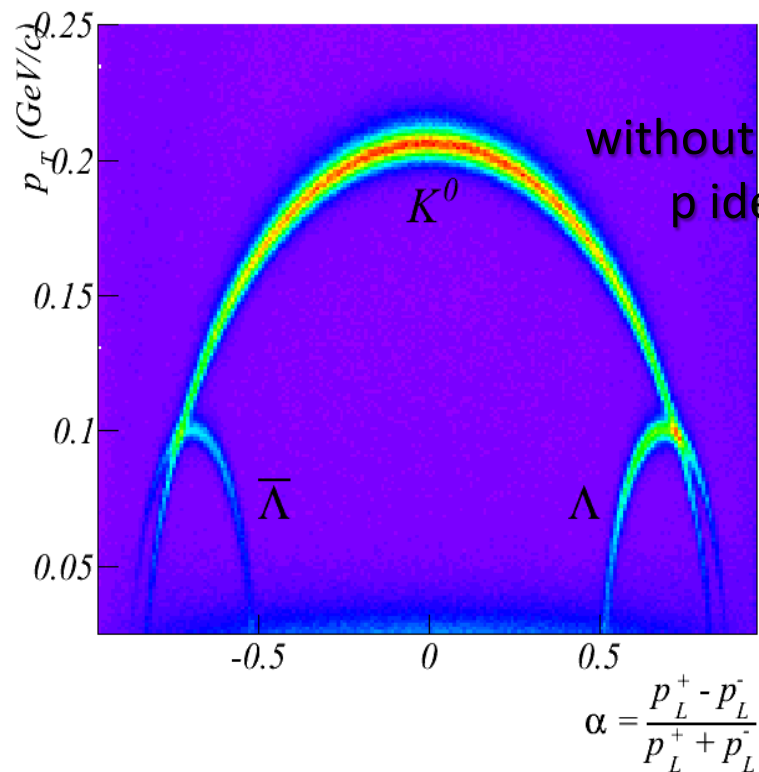
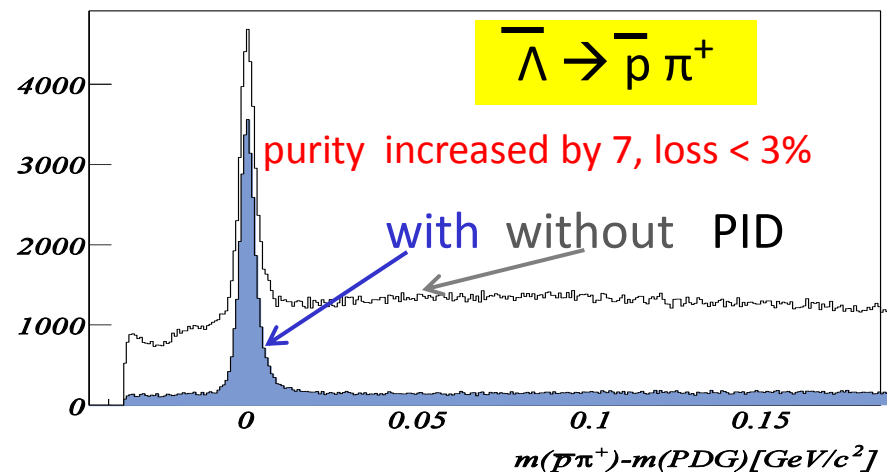
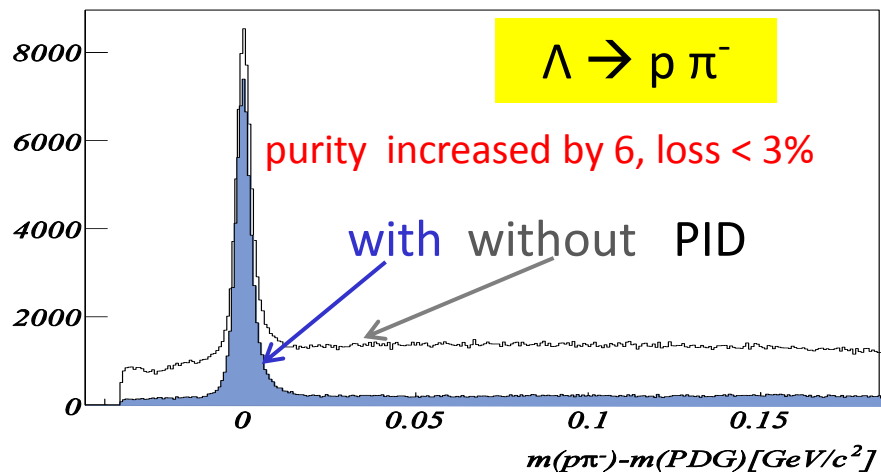
# The main components

- Al vessel ( $80 \text{ m}^3 \text{ C}_4\text{F}_{10}$ , 100 m of O-rings) with light windows
- Spherical UV mirrors ( $21 \text{ m}^2$ , split in two halves, 116 mirrors)
- Light beam pipe filled with He
- 16 fused silica windows ( $600 \text{ mm} \times 600 \text{ mm} \times 5 \text{ mm}$ )
- 8 MWPC's with CsI- coated photocathodes ( $4.3 \text{ m}^2$ )
- 62 k analog r/o channels: APV + GESSICA
- 576 MAPMT's with individual fused silica lens telescopes
- 9.2 k TDC channels: MAD + DREISAM + F1
- A complex gas circulation and filtering system

# RICH PID information in the $D^0$ analysis



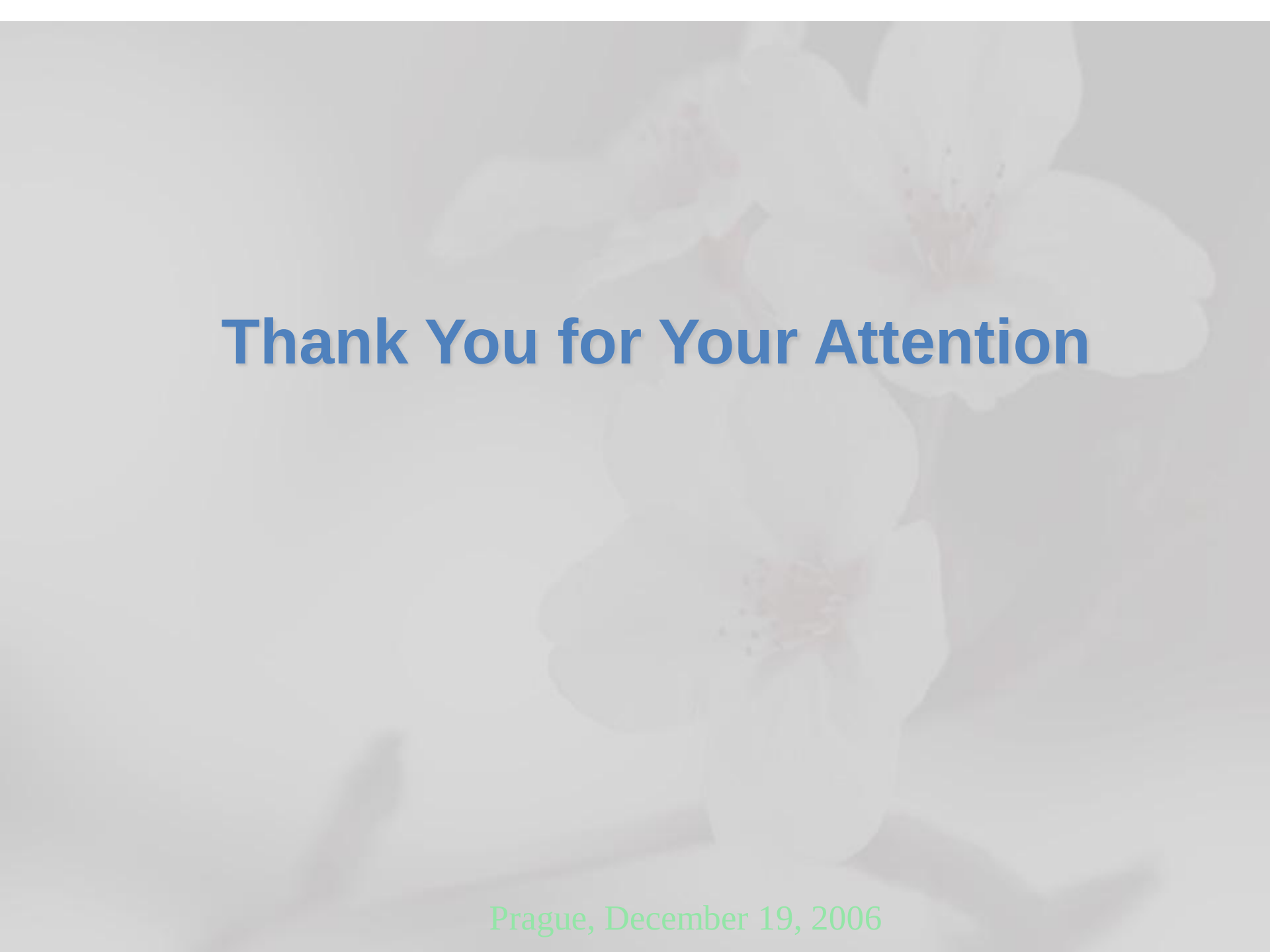
# RICH PID information in $\Lambda$ analysis



# The keys of the success

- A strong physics case
- A dedicated, highly motivated team
- A constant support from the Institutes and the Experiment

P. Abbon<sup>k</sup>, M. Alexeev<sup>a,1</sup>, H. Angerer<sup>i</sup>, R. Birsa<sup>o</sup>, P. Bordalo<sup>g,2</sup>, F. Bradamante<sup>n</sup>, A. Bressan<sup>n</sup>, M. Chiosso<sup>l</sup>, P. Ciliberti<sup>n</sup>, M. Colantoni<sup>m</sup>, T. Dafni<sup>k</sup>, S. Dalla Torre<sup>o,\*</sup>, E. Delagnes<sup>k</sup>, O. Denisov<sup>m</sup>, H. Deschamps<sup>k</sup>, V. Diaz<sup>o</sup>, N. Dibiase<sup>l</sup>, V. Duic<sup>n</sup>, W. Eyrich<sup>d</sup>, A. Ferrero<sup>l</sup>, M. Finger<sup>j</sup>, M. Finger Jr.<sup>j</sup>, H. Fischer<sup>e</sup>, S. Gerassimov<sup>i</sup>, M. Giorgi<sup>n</sup>, B. Gobbo<sup>o</sup>, R. Hagemann<sup>e</sup>, D. von Harrach<sup>h</sup>, F.H. Heinsius<sup>e</sup>, R. Joosten<sup>b</sup>, B. Ketzer<sup>i</sup>, V.N. Kolosov<sup>c,3</sup>, K. Königsman<sup>e</sup>, I. Konorov<sup>i</sup>, D. Kramer<sup>c,f</sup>, F. Kunne<sup>k</sup>, A. Lehmann<sup>d</sup>, S. Levorato<sup>n</sup>, A. Maggiora<sup>m</sup>, A. Magnon<sup>k</sup>, A. Mann<sup>i</sup>, A. Martin<sup>n</sup>, G. Menon<sup>o</sup>, A. Mutter<sup>e</sup>, O. Nähle<sup>b</sup>, F. Nerling<sup>e</sup>, D. Neyret<sup>k</sup>, D. Panzieri<sup>a</sup>, S. Paul<sup>i</sup>, G. Pesaro<sup>n</sup>, C. Pizzolotto<sup>d</sup>, J. Polak<sup>f,o</sup>, P. Rebougeard<sup>k</sup>, F. Robinet<sup>k</sup>, E. Rocco<sup>l</sup>, P. Schiavon<sup>n</sup>, C. Schill<sup>e</sup>, P. Schoenmeier<sup>d</sup>, W. Schröder<sup>d</sup>, L. Silva<sup>g</sup>, M. Slunecka<sup>j</sup>, F. Sozzi<sup>n</sup>, L. Steiger<sup>j</sup>, M. Sulc<sup>f</sup>, M. Svec<sup>f</sup>, S. Takekawa<sup>n</sup>, F. Tassarotto<sup>o</sup>, A. Teufel<sup>d</sup>, H. Wollny<sup>e</sup>



**Thank You for Your Attention**

Prague, December 19, 2006





